

# Lifetime measurements and oscillator strengths in singly ionized scandium and the solar abundance of scandium

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## ABSTRACT

The lifetimes of 17 even-parity levels (3d5s, 3d4d, 3d6s and 4p<sup>2</sup>) in the region 57 743–77 837 cm<sup>−1</sup> of singly ionized scandium (Sc II) were measured by two-step time-resolved laser induced fluorescence spectroscopy. Oscillator strengths of 57 lines from these highly excited upper levels were derived using a hollow cathode discharge lamp and a Fourier transform spectrometer. In addition, Hartree–Fock calculations where both the main relativistic and core-polarization effects were taken into account were carried out for both low- and high-excitation levels. There is a good agreement for most of the lines between our calculated branching fractions and the measurements of Lawler & Dakin in the region 9000–45 000 cm<sup>−1</sup> for low excitation levels and with our measurements for high excitation levels in the region 23 500–63 100 cm<sup>−1</sup>. This, in turn, allowed us to combine the calculated branching fractions with the available experimental lifetimes to determine semi-empirical oscillator strengths for a set of 380 E1 transitions in Sc II. These oscillator strengths include the weak lines that were used previously to derive the solar abundance of scandium. The solar abundance of scandium is now estimated to  $\log \epsilon_{\odot} = 3.04 \pm 0.13$  using these semi-empirical oscillator strengths to shift the values determined by Scott et al. The new estimated abundance value is in agreement with the meteoritic value ( $\log \epsilon_{\text{met}} = 3.05 \pm 0.02$ ) of Lodders, Palme & Gail.

**Key words:** atomic data – methods: laboratory: atomic – methods: numerical – techniques: spectroscopic – Sun: abundances.

## 1 INTRODUCTION

The iron-group elements ( $21 \leq Z \leq 28$ ) are produced during supernova type Ia explosions, while supernova type II explosions are responsible for the formation of  $\alpha$ -elements such as Mg, Si and S. The even- $Z$  nuclei such as S, Ca, Ti, Cr and Fe have higher cosmic abundance compared to the odd- $Z$  nuclei located in between because of the consecutive capture of  $\alpha$ -particles. The production of odd- $Z$  elements is not well understood and does not follow the abundance trends of the  $\alpha$ -elements, indicating non-common production mechanisms. In recent years, this has caused an increasing interest in the odd- $Z$  iron-peak elements in astrophysics. Abundance determinations in stars constrain the stellar evolution and supernova explosion models (Pagel 2009). Moreover, transitions from highly excited levels have an additional diagnostic value since they can be used to

benchmark non-local thermodynamical equilibrium (NLTE) modelling of stellar atmospheres. Besides the development of 3D hydrodynamic model atmospheres, a trustworthy NLTE treatment is the current challenge for accurate stellar abundances. High-precision atomic data for selected lines are important for this development (Lind, Bergmann & Asplund 2012).

In the case of scandium ( $Z = 21$ ), a realistic 3D NLTE solar atmosphere model has been used by Scott et al. (2015) to revise the solar abundance of scandium resulting in a photospheric value in significant disagreement with the meteoritic abundance (Lodders et al. 2009). Scott et al. (2015) used experimental transition probabilities of five Sc I and nine Sc II lines determined by Lawler & Dakin (1989). The latter authors combined their measured branching fractions with the time-resolved laser induced fluorescence (TR-LIF) lifetimes of Marsden et al. (1988) to obtain absolute  $A$ -values for transitions depopulating 51 levels in Sc I and 18 levels in Sc II. In Marsden et al. (1988), only three highly excited even-parity levels of Sc II, belonging to 3d4d <sup>3</sup>G, were measured. Older

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lifetime measurements in singly ionized scandium have focused on lower excited odd-parity  $3d4p$  and  $4s4p$  levels (Buchta et al. 1971; Arnesen et al. 1976; Palenius, Curtis & Lundlin 1976; Vogel et al. 1985). On the theoretical side, the most recent calculations of E1 oscillator strengths in Sc II are given in Ruczkowski, Elantkowska & Dembczynski (2014) and Kurucz (2011).

The main goal of this work is to provide a new set of experimental  $f$ -values for transitions depopulating the highly excited even-parity levels in Sc II, and new calculations for both low- and high-excitation levels and lines. Descriptions of our measurements are presented in Sections 2 and 3. The theoretical method used for the calculation of the radiative parameters is described in Section 4. In Section 5, our results are presented and compared to data available in the literature. The consequence of the proposed set of oscillator strengths on the solar abundance of scandium is discussed in Section 6. Finally, our conclusions are given in Section 7.

## 2 LIFETIME MEASUREMENTS

The experimental set-up for the two-step Time-Resolved Laser Induced Fluorescence (TR-LIF) measurements at the Lund High Power Laser Facility has been described in detail by Engström et al. (2014) and Lundberg et al. (2016). For an overview, we refer to fig. 1 in Lundberg et al. (2016), and here we give only the most important details. A frequency doubled Nd:YAG laser (Continuum Surelite) with 10 ns pulses was used to produce the free scandium ions by focusing the light on a rotating solid scandium sample in a vacuum chamber with a pressure of around  $10^{-4}$  mbar. The ions in the plasma cone were crossed by two laser beams, a few millimetre above the solid sample, generating the two-step excitations. The fluorescence signal was detected in a direction perpendicular to both the ablation and excitation lasers.

For the first step ( $4s-4p$ ), we used a Continuum Nd-60 dye laser with either DCM or Pyridine 2 dyes. The 10 ns long pulses were frequency doubled using a KDP crystal, giving the wavelengths needed for the first step. The second laser system excited the final high-energy levels. It consists of a frequency doubled Continuum NY-82 Nd:YAG laser pumping a Continuum Nd-60 dye laser with either DCM or Oxacin dye for wavelengths below or above 660 nm, respectively. The pulse length was reduced from 10 ns to less than 1 ns by stimulated Brillouin scattering. The output was frequency doubled using a KDP crystal and, where higher energy was needed, tripled with a BBO crystal.

For two-step excitation, the timing between the pulses is crucial. For this purpose, a delay generator ensures that the second step is timed to when the population of the intermediate state is at its flat maximum as determined by observing the decay of this level in another channel, see fig. 2 in Lundberg et al. (2016).

The fluorescence emitted by the scandium ions was filtered by a 1/8 m grating monochromator with its 0.28 mm wide entrance slit oriented parallel to the excitation laser beams. This fluorescence light was recorded using a fast micro-channel-plate photomultiplier tube (Hamamatsu R3809U) and digitized using a Tektronix DPO 7254 oscilloscope with 2.5 GHz analogue bandwidth. We used the second spectral order with a 0.5 nm observed line width for all measurements. The excitation laser pulse shape was recorded simultaneously using a fast photo diode and digitized by another channel of the oscilloscope. All decay curves were averaged over 1000 laser pulses and analysed using the `DECFIT` software (Palmeri et al. 2008) by fitting a single exponential function convoluted by

the measured shape of the second-step laser pulse and a background function to the observed decay.

The excitation schemes of the measured Sc II levels are presented in Table 1. This table shows the intermediate levels and their excitation wavelengths, the final levels and their excitation wavelengths from the intermediate levels together with the detection channel level and wavelength. For the levels  $4d\ ^3S_1$ ,  $4d\ ^1D_2$  and  $4p^2\ ^3P_2$ , it was possible to record the decay in more than one channel. We did not find any differences in the lifetimes obtained from the different channels. Sc II is a complex spectrum with a dense level structure, as shown in Fig. 1. Line blending can be caused by cascades or fluorescence from the intermediate level as discussed by Lundberg et al. (2016). For all measurements, we investigated if there was a line blend affecting the recorded curves. Due to the small spectral width of the laser compared to the energy level separations, we avoid exciting multiple levels.

To investigate any possible saturation effects in the second-step excitation, a set of neutral density filters was placed in the excitation beam. The delay between the ablation and first excitation pulse, the geometrical alignment of the lasers with respect to the target as well as the intensity of the ablation laser were varied to test time-of-flight effects. No systematic effects were observed.

As discussed in Palmeri et al. (2008), the weighting of individual data points, hence the purely statistical uncertainty in the fitted lifetime, is difficult to estimate accurately because the digitizing process is not strictly a counting measurement. However, extensive tests have shown that even for weak lines the dominating factor is the variation between different measurements. The uncertainty in Table 2 represents the uncertainty of 10–20 measurements performed over several days. The difference between subsequent curves is significantly lower than the quoted uncertainty, usually less than 1 per cent.

## 3 BRANCHING FRACTION MEASUREMENTS

A water-cooled hollow cathode discharge lamp (HCL) was used to produce the free scandium ions. The lamp has an iron cathode with anodes on each side, separated by glass cylinders. A small piece of scandium was placed in the cathode. We used argon, with a pressure of 0.3 Torr, as a buffer gas and applied currents ranging from 0.2 to 0.5 A. These measurements at different currents are very important to find and compensate for self-absorption effects. If self-absorption is not treated correctly, the measured relative line intensity may be less than the true intensity of the line. This in turn changes the branching fraction that is essential to derive oscillator strengths. Self-absorption was observed in the case of the  $3d4d\ ^3D_3$ ,  $3d4d\ ^3S_1$  and  $3d4d\ ^3P_2$  levels, and the affected lines were corrected. More details on this procedure can be found in Pehlivan, Nilsson & Hartman (2015).

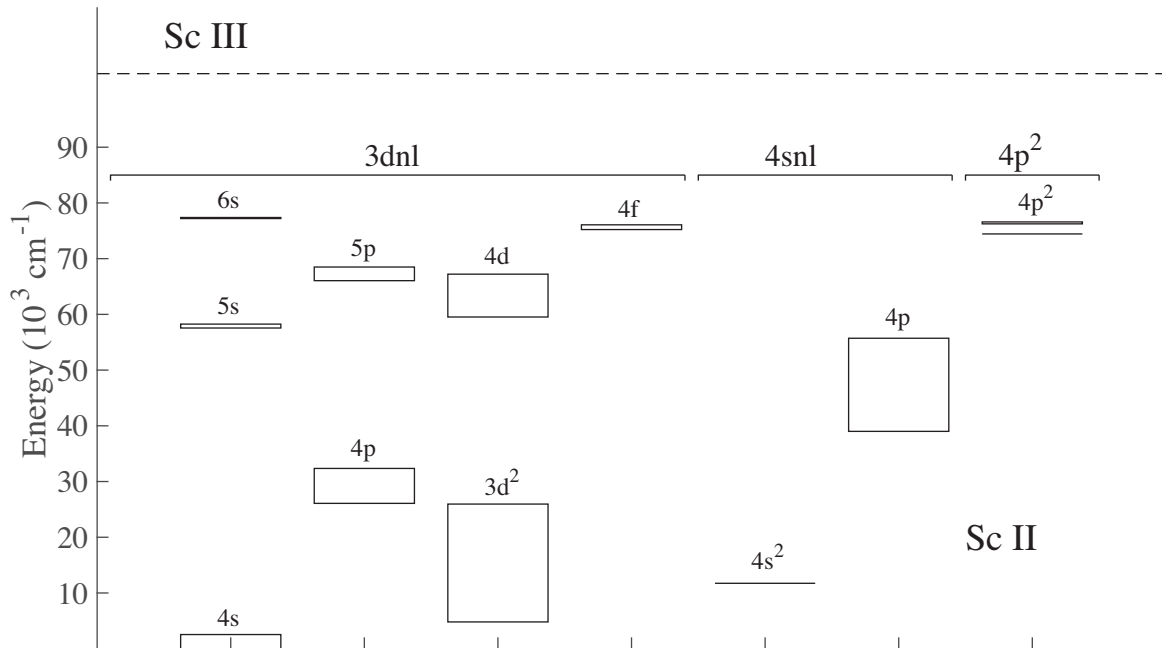
The spectra were recorded with the vacuum ultraviolet Fourier transform spectrometer (VUV FTS) at the Blackett Laboratory, Imperial College London (Pickering 2002) in the interval  $23\,500\text{--}63\,100\ \text{cm}^{-1}$  (425–158 nm) using a resolution of  $0.039\ \text{cm}^{-1}$ . We used two different photomultiplier tube detectors: Hamamatsu R7154 and R11568, the latter with a UG5 filter. Each scandium measurement consists of 12 co-added scans. To determine the relative response functions of the system, we used standard lamps: a tungsten filament lamp (800–300 nm) and a deuterium lamp (410–116 nm) for the wavelength region (425–210 nm), and a deuterium standard lamp alone for the region (317–158 nm). The tungsten lamp was calibrated by the UK National Physical Laboratory

**Table 1.** Measured Sc II levels and the corresponding two-step excitation schemes.

Final level <sup>a</sup>	First-step excitation			Second-step excitation			Detection	
	Starting level <sup>a</sup> (cm <sup>-1</sup> )	Intermediate level <sup>a</sup> (cm <sup>-1</sup> )	$\lambda_{\text{air}}$ (nm)	Final level <sup>a</sup> (cm <sup>-1</sup> )	$\lambda_{\text{air}}$ (nm)	Scheme <sup>b</sup>	Lower level <sup>a</sup> (cm <sup>-1</sup> )	$\lambda_{\text{air}}$ (nm)
5s <sup>3</sup> D <sub>3</sub>	67.72	27 602.45	363.07	57 743.92	331.67	2 $\omega$	27 841.35	334.32
5s <sup>1</sup> D <sub>2</sub>	67.72	27 602.45	363.07	58 252.09	326.17	2 $\omega$	32 349.98	385.96
4d <sup>1</sup> F <sub>3</sub>	177.76	29 823.93	337.22	59 528.42	336.55	2 $\omega$	26 081.34	298.89
4d <sup>3</sup> D <sub>1</sub>	177.76	29 823.93	337.22	59 875.08	332.67	2 $\omega$	27 917.78	312.83
4d <sup>3</sup> D <sub>2</sub>	177.76	29 823.93	337.22	59 929.46	332.07	2 $\omega$	28 021.29	313.31
4d <sup>3</sup> D <sub>3</sub>	177.76	29 823.93	337.22	60 001.91	331.27	2 $\omega$	28 161.17	313.97
4d <sup>3</sup> G <sub>3</sub>	177.76	29 823.93	337.22	60 267.16	328.39	2 $\omega$	27 443.71	304.57
4d <sup>1</sup> P <sub>1</sub>	177.76	29 823.93	337.22	60 400.41	326.95	2 $\omega$	26 081.34	291.30
4d <sup>3</sup> S <sub>1</sub>	177.76	29 823.93	337.22	61 071.43	319.93	2 $\omega$	29 823.93	319.93
							39 345.52	460.15
4d <sup>3</sup> F <sub>2</sub>	2540.95	32 349.98	335.37	63 374.63	322.23	2 $\omega$	27 917.78	281.95
4d <sup>3</sup> F <sub>4</sub>	2540.95	32 349.98	335.37	63 528.54	320.64	2 $\omega$	28 161.17	282.66
4d <sup>1</sup> D <sub>2</sub>	2540.95	32 349.98	335.37	64 366.68	312.25	2 $\omega$	26 081.34	261.12
							30 815.70	297.97
4d <sup>3</sup> P <sub>2</sub>	2540.95	32 349.98	335.37	64 705.89	308.98	2 $\omega$	29 823.93	286.60
4p <sup>2</sup> <sup>1</sup> D <sub>2</sub>	177.76	28 161.17	357.25	74 433.30	216.04	3 $\omega$	32 349.98	237.55
4p <sup>2</sup> <sup>3</sup> P <sub>1</sub>	177.76	29 823.93	337.22	76 360.80	214.82	3 $\omega$	39 345.52	270.08
4p <sup>2</sup> <sup>3</sup> P <sub>2</sub>	177.76	29 823.93	337.22	76 589.30	213.76	3 $\omega$	28 161.17	206.43
							39 115.04	266.77
							39 345.52	268.42
6s <sup>3</sup> D <sub>3</sub>	177.76	29 823.93	337.22	77 387.17	210.18	3 $\omega$	28 161.17	203.08

Notes. <sup>a</sup>All energy-level values and wavelength values are from Johansson & Litzén (1980).

<sup>b</sup>2 $\omega$  and 3 $\omega$  stand for, respectively, frequency doubling and tripling excitation schemes. All first-step levels are excited using a frequency doubling scheme (2 $\omega$ ).


**Figure 1.** Partial energy level diagram of Sc II, the energy level values are from Johansson & Litzén (1980). Each box consists of several levels.

and the deuterium lamp by Physikalisch-Technische Bundesanstalt, in Berlin. In the region where the lamps overlap, the response functions were placed on the same relative scale. We recorded the spectrum of the calibration lamps immediately before and after each scandium measurement. The HCL and the calibration lamps were placed at the same distances from the FTS, and a mirror was used to select the light source without moving the lamps.

In astrophysics, oscillator strengths ( $f$ -values) or  $\log(gf)$  values are the parameters used for abundance analysis. The  $f$ -value is proportional to the transition probability for E1 transitions by

$$f = \frac{g_u}{g_l} \lambda^2 A_{ul} 1.499 \times 10^{-16}, \quad (1)$$

**Table 2.** A comparison of radiative lifetimes ( $\tau$ ) in Sc II.

Level <sup>a</sup>	$E^a$ (cm <sup>-1</sup> )	$\tau_{\text{this cal}}^b$ (ns)	$\tau_{\text{this exp}}^c$ (ns)	$\tau_{\text{other exp}}$ (ns)	$\tau_{\text{other cal}}$ (ns)
3d4p <sup>1</sup> D <sub>2</sub> <sup>o</sup>	26 081.34	6.65		7.5 ± 0.4 <sup>d</sup> 7.16 ± 0.18 <sup>e</sup>	6.54 <sup>i</sup> 7.79 <sup>j</sup>
3d4p <sup>3</sup> F <sub>2</sub> <sup>o</sup>	27 443.71	5.68		6.2 ± 0.3 <sup>d</sup> 6.2 ± 0.2 <sup>f</sup> 6.5 ± 0.4 <sup>g</sup>	5.38 <sup>i</sup> 5.90 <sup>j</sup>
3d4p <sup>3</sup> F <sub>3</sub> <sup>o</sup>	27 602.45	5.62		6.1 ± 0.3 <sup>d</sup>	5.32 <sup>i</sup> 5.83 <sup>j</sup>
3d4p <sup>3</sup> F <sub>4</sub> <sup>o</sup>	27 841.35	5.54		6.1 ± 0.3 <sup>d</sup> 5.6 ± 0.6 <sup>h</sup>	5.24 <sup>i</sup> 5.75 <sup>j</sup>
3d4p <sup>3</sup> D <sub>1</sub> <sup>o</sup>	27 917.78	4.44		4.7 ± 0.2 <sup>d</sup> 4.61 ± 0.10 <sup>e</sup>	4.20 <sup>i</sup> 4.67 <sup>j</sup>
3d4p <sup>3</sup> D <sub>2</sub> <sup>o</sup>	28 021.29	4.41		4.7 ± 0.2 <sup>d</sup> 4.66 ± 0.14 <sup>e</sup>	4.17 <sup>i</sup> 4.64 <sup>j</sup>
3d4p <sup>3</sup> D <sub>3</sub> <sup>o</sup>	28 161.17	4.38		4.7 ± 0.2 <sup>d</sup> 4.55 ± 0.15 <sup>e</sup> 6.1 ± 0.6 <sup>h</sup>	4.15 <sup>i</sup> 4.59 <sup>j</sup>
3d4p <sup>3</sup> P <sub>0</sub> <sup>o</sup>	29 736.27	6.36		7.7 ± 0.4 <sup>d</sup> 7.48 ± 0.18 <sup>e</sup>	6.80 <sup>i</sup> 7.44 <sup>j</sup>
3d4p <sup>3</sup> P <sub>1</sub> <sup>o</sup>	29 742.16	6.39		7.6 ± 0.4 <sup>d</sup> 7.3 ± 0.3 <sup>e</sup>	6.76 <sup>i</sup> 7.45 <sup>j</sup>
3d4p <sup>3</sup> P <sub>2</sub> <sup>o</sup>	29 823.93	6.30		7.4 ± 0.4 <sup>d</sup> 7.30 ± 0.16 <sup>e</sup>	6.67 <sup>i</sup> 7.50 <sup>j</sup>
3d4p <sup>1</sup> P <sub>1</sub> <sup>o</sup>	30 815.70	8.10		8.8 ± 0.4 <sup>d</sup> 8.5 ± 0.6 <sup>g</sup> 5.5 ± 0.5 <sup>h</sup>	7.35 <sup>i</sup> 8.76 <sup>j</sup>
3d4p <sup>1</sup> F <sub>3</sub> <sup>o</sup>	32 349.98	4.68		5.1 ± 0.3 <sup>d</sup> 5.2 ± 0.2 <sup>e</sup> 6.8 ± 0.6 <sup>h</sup>	4.46 <sup>i</sup> 5.20 <sup>j</sup>
4s4p <sup>3</sup> P <sub>0</sub> <sup>o</sup>	39 002.20	3.69		3.7 ± 0.2 <sup>d</sup>	3.36 <sup>i</sup> 3.66 <sup>j</sup>
4s4p <sup>3</sup> P <sub>1</sub> <sup>o</sup>	39 115.04	3.69		3.7 ± 0.2 <sup>d</sup>	3.37 <sup>i</sup> 3.67 <sup>j</sup>
4s4p <sup>3</sup> P <sub>2</sub> <sup>o</sup>	39 345.52	3.70		3.8 ± 0.2 <sup>d</sup>	3.39 <sup>i</sup> 3.67 <sup>j</sup>
4s4p <sup>1</sup> P <sub>1</sub> <sup>o</sup>	55 715.36	0.88			0.91 <sup>i</sup>
3d5s <sup>3</sup> D <sub>1</sub>	57 551.88	3.49			3.44 <sup>i</sup>
3d5s <sup>3</sup> D <sub>2</sub>	57 614.40	3.50			3.44 <sup>i</sup>
3d5s <sup>3</sup> D <sub>3</sub>	57 743.92	3.50	3.20 ± 0.20		3.44 <sup>i</sup>
3d5s <sup>1</sup> D <sub>2</sub>	58 252.09	3.70	3.26 ± 0.20		3.66 <sup>j</sup>
3d4d <sup>1</sup> F <sub>3</sub>	59 528.42	2.69	2.32 ± 0.15		2.51 <sup>i</sup> 2.51 <sup>j</sup>
3d4d <sup>3</sup> D <sub>1</sub>	59 875.08	2.72	2.23 ± 0.15		2.62 <sup>i</sup>
3d4d <sup>3</sup> D <sub>2</sub>	59 929.46	2.74	2.32 ± 0.15		2.63 <sup>i</sup> 2.58 <sup>j</sup>
3d4d <sup>3</sup> D <sub>3</sub>	60 001.91	2.76	2.41 ± 0.20		2.65 <sup>i</sup>
3d4d <sup>3</sup> G <sub>3</sub>	60 267.16	2.50	2.19 ± 0.15	2.5 ± 0.2 <sup>d</sup>	2.33 <sup>i</sup> 2.47 <sup>j</sup>
3d4d <sup>3</sup> G <sub>4</sub>	60 348.46	2.52		2.4 ± 0.2 <sup>d</sup>	2.35 <sup>i</sup> 2.49 <sup>j</sup>
3d4d <sup>1</sup> P <sub>1</sub>	60 400.41	2.89	2.44 ± 0.15		2.69 <sup>j</sup> 2.63 <sup>j</sup>
3d4d <sup>3</sup> G <sub>5</sub>	60 457.12	2.54		2.5 ± 0.2 <sup>d</sup>	2.38 <sup>i</sup> 2.51 <sup>j</sup>
3d4d <sup>3</sup> S <sub>1</sub>	61 071.43	2.82	2.45 ± 0.15		2.77 <sup>i</sup> 2.78 <sup>j</sup>
3d4d <sup>3</sup> F <sub>2</sub>	63 374.63	2.40	2.15 ± 0.10		2.05 <sup>i</sup> 2.43 <sup>j</sup>
3d4d <sup>3</sup> F <sub>3</sub>	63 445.16	2.41			2.05 <sup>i</sup>
3d4d <sup>3</sup> F <sub>4</sub>	63 528.54	2.43	2.19 ± 0.10		2.07 <sup>i</sup> 2.47 <sup>j</sup>
3d4d <sup>1</sup> D <sub>2</sub>	64 366.68	2.73	2.25 ± 0.15		2.26 <sup>i</sup>

**Table 2** – *continued*

Level <sup>a</sup>	$E^a$ (cm <sup>-1</sup> )	$\tau_{\text{this cal}}^b$ (ns)	$\tau_{\text{this exp}}^c$ (ns)	$\tau_{\text{other exp}}$ (ns)	$\tau_{\text{other cal}}$ (ns)
3d4d <sup>3</sup> P <sub>0</sub>	64 615.77	3.21			2.65 <sup>i</sup>
3d4d <sup>3</sup> P <sub>1</sub>	64 646.70	3.21			2.65 <sup>i</sup>
3d4d <sup>3</sup> P <sub>2</sub>	64 705.89	3.19	2.51 ± 0.15		2.65 <sup>i</sup>
3d4d <sup>1</sup> G <sub>4</sub>	65 236.04	3.17			2.45 <sup>i</sup>
3d4d <sup>1</sup> S <sub>0</sub>	67 216.56	3.87			2.74 <sup>i</sup>
4p <sup>2</sup> <sup>1</sup> D <sub>2</sub>	74 433.30	5.96	3.80 ± 0.15		6.80 <sup>i</sup>
4p <sup>2</sup> <sup>3</sup> P <sub>0</sub>	76 243.20	1.17			1.28 <sup>i</sup>
4p <sup>2</sup> <sup>3</sup> P <sub>1</sub>	76 360.80	1.17	1.14 ± 0.06		1.28 <sup>i</sup>
4p <sup>2</sup> <sup>3</sup> P <sub>2</sub>	76 589.30	1.18	1.09 ± 0.06		1.30 <sup>i</sup> 1.03 <sup>j</sup>
3d6s <sup>3</sup> D <sub>1</sub>	77 195.19	5.56			5.00 <sup>i</sup>
3d6s <sup>3</sup> D <sub>2</sub>	77 256.99	5.55			5.00 <sup>i</sup>
3d6s <sup>3</sup> D <sub>3</sub>	77 387.17	5.54	3.73 ± 0.25		4.98 <sup>i</sup>
3d6s <sup>1</sup> D <sub>2</sub>	77 833.88	6.61			6.94 <sup>i</sup>

Notes. <sup>a</sup>Johansson & Litzén (1980).

<sup>b</sup>HFR+CPOL calculation, this work.

<sup>c</sup>TR-LIF measurements, this work.

<sup>d</sup>TR-LIF measurements by Marsden et al. (1988).

<sup>e</sup>TR-LIF measurements by Vogel et al. (1985).

<sup>f</sup>TR-LIF measurements by Arnesen et al. (1976).

<sup>g</sup>Beam-foil spectroscopy by Palenius et al. (1976).

<sup>h</sup>Beam-foil spectroscopy by Buchta et al. (1971).

<sup>i</sup>HFR calculation by Kurucz (2011).

<sup>j</sup>Parametric method calculation by Ruczkowski et al. (2014).

where  $g_u$  is the statistical weight of the upper level,  $g_l$  the statistical weight of the lower level,  $\lambda$  the wavelength of the transition in Å and  $A_{ul}$  the transition probability in s<sup>-1</sup>.

The transition probability is related to the branching fraction ( $BF$ ) and the lifetime of the upper level ( $\tau_u$ ). It can be derived using

$$A_{ul} = \frac{BF_{ul}}{\tau_u}. \quad (2)$$

We obtained the lifetimes of the upper levels from our measurements, as discussed in Section 2. The  $BF$  is the parameter we measure and it is defined as the transition probability of the line,  $A_{ul}$ , divided by the sum of transition probabilities of all lines from the same upper level;

$$BF_{ul} = \frac{A_{ul}}{\sum_i A_{ui}} = \frac{I_{ul}}{\sum_i I_{ui}}. \quad (3)$$

Since all lines emanate from the same upper level, the transition probability is proportional to the line intensity,  $I_{ul}$ , which for FTS spectra is proportional to photon flux (Davis, Abrams & Brault 2001). Therefore, we derived  $BF$ s from calibrated intensity ratios in our measurements. All lines were identified using the analysis of Johansson & Litzén (1980). The intensities of the observed lines were determined by fitting Gaussian line profiles using GFIT (Engström 1998, 2014).

The uncertainty of the  $A$ -value, and thus of the  $f$ -value, arises from the uncertainty in the upper level lifetime and the uncertainty of the  $BF$ . The latter includes the uncertainty in the intensity calibration procedure and the uncertainty in the measured line intensity, including the self-absorption correction. The uncertainties of the integrated line intensities were determined using GFIT. The relative uncertainties are as low as 0.1 per cent for strong lines and 4 per cent on average. However, for two weak lines the uncertainty is as large as 20 per cent. The uncertainty in the calibration using the tungsten lamp is 2.2 per cent and the uncertainty using the deuterium lamp is 8.6 per cent for the region 425–210 nm and 9.9 per cent for

**Table 3.** Presentation of experimental  $\log(gf)$  values together with the transition, wavelength,  $\lambda$ , wavenumber,  $\sigma$ , measured branching fraction,  $BF_{\text{exp}}$ , experimental transition probability,  $A_{\text{exp}}$  and the corresponding rescaled semi-empirical  $\log(gf)$  values of this work. The radiative lifetimes,  $\tau$ , are TR-LIF measurements from this work.

Config.	Upper level <sup>a</sup>		Lower level <sup>a</sup>		$\lambda_{\text{exp}}^a$ (nm)	$\sigma_{\text{exp}}^a$ (cm <sup>-1</sup> )	$\sigma_{\text{theo}}^b$ (cm <sup>-1</sup> )	$BF_{\text{exp}}$	$BF$ unc. per cent	$A_{\text{exp}}$ (s <sup>-1</sup> )	$\log(gf)$ Exp.	$\log(gf)_{\text{resc}}$ Calc.
	Energy (cm <sup>-1</sup> )	Config.	Energy (cm <sup>-1</sup> )	Config.								
3d5s <sup>3</sup> D <sub>3</sub> $\tau = 3.20 \pm 0.20$ ns	57 744	3d4p	<sup>3</sup> F <sub>3</sub> <sup>o</sup>	27 602	331.673	30 141.50	30 176	6.21E-02	4	1.94E+07	-0.65 ± 0.03	-0.75
			<sup>3</sup> F <sub>4</sub> <sup>o</sup>	27 841	334.323	29 902.57	29 944	4.05E-01	3	1.27E+08	0.17 ± 0.03	0.14
			<sup>3</sup> D <sub>2</sub> <sup>o</sup>	28 021	336.347	29 722.58	29 771	5.29E-02	4	1.65E+07	-0.71 ± 0.03	-0.69
			<sup>3</sup> D <sub>3</sub> <sup>o</sup>	28 161	337.938	29 582.76	29 620	3.60E-01	3	1.12E+08	0.13 ± 0.03	0.17
			<sup>3</sup> P <sub>2</sub> <sup>o</sup>	29 824	358.064	27 919.88	27 888	1.20E-01	4	3.75E+07	-0.30 ± 0.03	-0.28
			<i>Residual</i>							3.37E-03		
3d5s <sup>1</sup> D <sub>2</sub> $\tau = 3.26 \pm 0.20$ ns	58 252	3d4p	<sup>1</sup> D <sub>2</sub> <sup>o</sup>	26 081	310.751	32 179.68	32 040	4.90E-01	2	1.50E+08	0.04 ± 0.03	-0.09
			<sup>3</sup> F <sub>2</sub> <sup>o</sup>	27 444	324.493	30 808.34	30 866	1.08E-02	16	3.30E+06	-1.58 ± 0.07	-1.64
			<sup>1</sup> P <sub>1</sub> <sup>o</sup>	30 816	364.376	27 436.43	27 508	1.48E-01	4	4.55E+07	-0.34 ± 0.03	-0.32
			<sup>1</sup> F <sub>3</sub> <sup>o</sup>	32 350	385.960	25 902.13	25 865	3.51E-01	5	1.08E+08	0.08 ± 0.03	0.18
			<i>Residual</i>							2.10E-02		
3d4d <sup>1</sup> F <sub>3</sub> $\tau = 2.32 \pm 0.15$ ns	59 528	3d4p	<sup>1</sup> D <sub>2</sub> <sup>o</sup>	26 081	298.893	33 447.17	33 296	8.22E-01	0.5	3.54E+08	0.52 ± 0.03	0.47
			<sup>3</sup> D <sub>3</sub> <sup>o</sup>	28 161	318.712	31 367.21	31 408	6.54E-03	14	2.82E+06	-1.52 ± 0.06	-1.45
			<sup>1</sup> F <sub>3</sub> <sup>o</sup>	32 350	367.834	27 178.50	27 121	1.72E-01	7	7.40E+07	0.02 ± 0.04	0.20
			<i>Residual</i>							5.10E-03		
3d4d <sup>3</sup> D <sub>1</sub> $\tau = 2.23 \pm 0.15$ ns	59 875	3d4p	<sup>3</sup> F <sub>2</sub> <sup>o</sup>	27 444	308.254	32 431.14	32 475	1.22E-01	5	5.49E+07	-0.63 ± 0.04	-0.63
			<sup>3</sup> D <sub>1</sub> <sup>o</sup>	27 918	312.827	31 957.28	32 026	4.57E-01	3	2.05E+08	-0.04 ± 0.03	-0.11
			<sup>3</sup> D <sub>2</sub> <sup>o</sup>	28 021	313.843	31 853.76	31 913	1.17E-01	5	5.24E+07	-0.63 ± 0.04	-0.65
			<sup>3</sup> P <sub>0</sub> <sup>o</sup>	29 736	331.703	30 138.84	30 134	1.79E-01	4	8.02E+07	-0.40 ± 0.03	-0.37
			<sup>3</sup> P <sub>1</sub> <sup>o</sup>	29 742	331.768	30 132.91	30 123	1.25E-01	5	5.61E+07	-0.56 ± 0.04	-0.50
			<i>Residual</i>							3.75E-02		
3d4d <sup>3</sup> D <sub>2</sub> $\tau = 2.32 \pm 0.15$ ns	59 929	3d4p	<sup>3</sup> F <sub>3</sub> <sup>o</sup>	27 602	309.249	32 327.05	32 372	1.21E-01	5	5.23E+07	-0.43 ± 0.03	-0.45
			<sup>3</sup> D <sub>1</sub> <sup>o</sup>	27 918	312.296	32 011.74	32 081	8.39E-02	5	3.62E+07	-0.58 ± 0.03	-0.63
			<sup>3</sup> D <sub>2</sub> <sup>o</sup>	28 021	313.309	31 908.30	31 968	4.12E-01	3	1.78E+08	0.12 ± 0.03	0.06
			<sup>3</sup> D <sub>3</sub> <sup>o</sup>	28 161	314.688	31 768.28	31 816	6.61E-02	5	2.85E+07	-0.67 ± 0.03	-0.68
			<sup>3</sup> P <sub>1</sub> <sup>o</sup>	29 742	331.170	30 187.30	30 178	2.50E-01	4	1.08E+08	-0.05 ± 0.03	-0.02
			<sup>3</sup> P <sub>2</sub> <sup>o</sup>	29 824	332.069	30 105.53	30 084	6.67E-02	5	2.87E+07	-0.62 ± 0.03	-0.56
			<i>Residual</i>							3.42E-02		
3d4d <sup>3</sup> D <sub>3</sub> $\tau = 2.41 \pm 0.20$ ns	60 002	3d4p	<sup>3</sup> F <sub>4</sub> <sup>o</sup>	27 841	310.850	32 160.62	32 214	9.39E-02	6	3.90E+07	-0.40 ± 0.04	-0.34
			<sup>3</sup> D <sub>2</sub> <sup>o</sup>	28 021	312.599	31 980.37	32 041	4.35E-02	7	1.81E+07	-0.73 ± 0.06	-0.63
			<sup>3</sup> D <sub>3</sub> <sup>o</sup>	28 161	313.972	31 840.77	31 890	5.16E-01	4	2.14E+08	0.35 ± 0.04	0.28
			<sup>3</sup> P <sub>2</sub> <sup>o</sup>	29 824	331.272	30 178.03	30 157	3.47E-01	5	1.44E+08	0.22 ± 0.04	0.24
			<i>Residual</i>							3.51E-02		
3d4d <sup>3</sup> G <sub>3</sub> $\tau = 2.19 \pm 0.15$ ns	60267	3d4p	<sup>3</sup> F <sub>2</sub> <sup>o</sup>	27 444	304.572	32 823.36	32 822	9.26E-01	1	4.23E+08	0.61 ± 0.03	0.61
			<sup>3</sup> F <sub>3</sub> <sup>o</sup>	27 602	306.052	32 664.51	32 664	7.41E-02	8	3.38E+07	-0.48 ± 0.04	-0.47
			<i>Residual</i>							6.20E-03		
3d4d <sup>1</sup> P <sub>1</sub> $\tau = 2.44 \pm 0.15$ ns	60400	3d4p	<sup>1</sup> D <sub>2</sub> <sup>o</sup>	26 081	291.298	34 319.09	34 206	3.98E-01	7	1.63E+08	-0.21 ± 0.04	-0.33
			<sup>1</sup> P <sub>1</sub> <sup>o</sup>	30 816	337.915	29 584.65	29 673	6.02E-01	5	2.47E+08	0.10 ± 0.03	0.14
			<i>Residual</i>							7.27E-02		
3d4d <sup>3</sup> S <sub>1</sub> $\tau = 2.45 \pm 0.15$ ns	61071	3d4p	<sup>3</sup> P <sub>0</sub> <sup>o</sup>	29 736	319.038	31 335.12	31 336	1.13E-01	7	4.60E+07	-0.68 ± 0.04	-0.70
			<sup>3</sup> P <sub>1</sub> <sup>o</sup>	29 742	319.098	31 329.24	31 326	2.84E-01	6	1.16E+08	-0.28 ± 0.04	-0.25
			<sup>3</sup> P <sub>2</sub> <sup>o</sup>	29 824	319.933	31 247.50	31 231	5.73E-01	4	2.34E+08	0.03 ± 0.03	-0.02
			<sup>1</sup> P <sub>1</sub> <sup>o</sup>	30 816	330.421	30 255.76	30 319	3.04E-02	10	1.24E+07	-1.21 ± 0.05	-1.28
			<i>Residual</i>							6.10E-02		
3d4d <sup>3</sup> F <sub>2</sub> $\tau = 2.15 \pm 0.10$ ns	63375	3d4p	<sup>3</sup> F <sub>2</sub> <sup>o</sup>	27 444	278.230	35 930.81	35 960	3.57E-01	5	1.66E+08	-0.02 ± 0.03	-0.07
			<sup>3</sup> F <sub>3</sub> <sup>o</sup>	27 602	279.464	35 772.19	35 802	3.72E-02	9	1.73E+07	-0.99 ± 0.04	-0.95
			<sup>3</sup> D <sub>1</sub> <sup>o</sup>	27 918	281.950	35 456.96	35 511	5.27E-01	3	2.45E+08	0.17 ± 0.02	0.17
			<sup>3</sup> D <sub>2</sub> <sup>o</sup>	28 021	282.776	35 353.30	35 398	7.88E-02	6	3.66E+07	-0.66 ± 0.03	-0.60
			<i>Residual</i>							1.59E-02		
3d4d <sup>3</sup> F <sub>4</sub> $\tau = 2.19 \pm 0.10$ ns	63529	3d4p	<sup>3</sup> F <sub>4</sub> <sup>o</sup>	27 841	280.130	35 687.12	35 726	3.42E-01	6	1.56E+08	0.22 ± 0.03	0.21
			<sup>3</sup> D <sub>3</sub> <sup>o</sup>	28 161	282.663	35 367.30	35 402	6.58E-01	3	3.01E+08	0.51 ± 0.02	0.51
			<i>Residual</i>							8.18E-03		
3d4d <sup>1</sup> D <sub>2</sub> $\tau = 2.25 \pm 0.15$ ns	64367	3d4p	<sup>1</sup> D <sub>2</sub> <sup>o</sup>	26 081	261.119	38 285.22	38 187	7.25E-01	4	3.22E+08	0.22 ± 0.03	0.11
			<sup>3</sup> F <sub>2</sub> <sup>o</sup>	27 444	270.754	36 923.00	37 012	1.69E-02	12	7.51E+06	-1.38 ± 0.06	-1.49
			<sup>3</sup> P <sub>1</sub> <sup>o</sup>	29 742	288.728	34 624.48	34 661	2.27E-02	16	1.01E+07	-1.20 ± 0.07	-1.07

**Table 3** – *continued*

Config.	Upper level <sup>a</sup> Energy (cm <sup>-1</sup> )	Lower level <sup>a</sup> Config. Energy (cm <sup>-1</sup> )	$\lambda_{\text{exp}}^a$ (nm)	$\sigma_{\text{exp}}^a$ (cm <sup>-1</sup> )	$\sigma_{\text{theo}}^b$ (cm <sup>-1</sup> )	$BF_{\text{exp}}$	$BF$ unc. per cent	$A_{\text{exp}}$ (s <sup>-1</sup> )	log (gf) Exp.	log (gf) <sub>resc</sub> Calc.	
		3d4p <sup>1</sup> P <sub>1</sub> <sup>o</sup> <i>Residual</i>	30 816	297.967	33 550.90	33 654	2.35E-01 4.94E-02	8	1.05E+08	-0.16 ± 0.04	-0.02
3d4d <sup>3</sup> P <sub>2</sub> $\tau = 2.51 \pm 0.15$ ns	64706	3d4p <sup>3</sup> D <sub>3</sub> <sup>o</sup>	28 161	273.556	36 544.66	36 597	1.59E-01	8	6.32E+07	-0.45 ± 0.04	-0.46
		3d4p <sup>3</sup> P <sub>1</sub> <sup>o</sup>	29 742	285.927	34 963.68	34 960	1.78E-01	6	7.09E+07	-0.36 ± 0.04	-0.40
		3d4p <sup>3</sup> P <sub>2</sub> <sup>o</sup>	29 824	286.597	34 881.86	34 865	6.33E-01	3	2.52E+08	0.19 ± 0.03	0.16
		3d4p <sup>1</sup> P <sub>1</sub> <sup>o</sup> <i>Residual</i>	30 816	294.984	33 890.19	33 953	3.08E-02 4.86E-02	8	1.23E+07	-1.10 ± 0.04	-0.98
4p <sup>2</sup> <sup>3</sup> P <sub>1</sub> $\tau = 1.14 \pm 0.06$ ns	76361	3d4p <sup>3</sup> D <sub>2</sub> <sup>o</sup>	28 021	206.804	48 339.50	48 384	3.16E-01	7	2.77E+08	-0.27 ± 0.04	-0.41
		4s4p <sup>3</sup> P <sub>0</sub> <sup>o</sup>	39 002	267.597	37 358.69	37 358	2.29E-01	6	2.01E+08	-0.19 ± 0.03	-0.18
		4s4p <sup>3</sup> P <sub>1</sub> <sup>o</sup>	39 115	268.407	37 245.53	37 245	1.80E-01	6	1.58E+08	-0.29 ± 0.03	-0.31
		4s4p <sup>3</sup> P <sub>2</sub> <sup>o</sup> <i>Residual</i>	39 346	270.079	37 014.70	37 014	2.75E-01 6.46E-02	6	2.41E+08	-0.10 ± 0.03	-0.09
4p <sup>2</sup> <sup>3</sup> P <sub>2</sub> $\tau = 1.09 \pm 0.06$ ns	76589	3d4p <sup>3</sup> D <sub>2</sub> <sup>o</sup>	28 021	205.831	48 568.03	48 615	5.63E-02	27	5.17E+07	-0.78 ± 0.11	-0.88
		3d4p <sup>3</sup> D <sub>3</sub> <sup>o</sup>	28 161	206.426	48 428.15	48 464	2.95E-01	6	2.71E+08	-0.06 ± 0.03	-0.13
		4s4p <sup>3</sup> P <sub>1</sub> <sup>o</sup>	39 115	266.771	37 474.35	37 477	1.65E-01	6	1.52E+08	-0.09 ± 0.03	-0.06
		4s4p <sup>3</sup> P <sub>2</sub> <sup>o</sup> <i>Residual</i>	39 346	268.422	37 243.72	37 246	4.83E-01 1.18E-02	4	4.43E+08	0.38 ± 0.03	0.41

Note. <sup>a</sup>Energy level, wavelength and wavenumber values are taken from Johansson & Litzén (1980) which are available in NIST data base (Kramida et al. 2015). <sup>b</sup>Theoretical wavenumber values are from the calculations of this work.

317–158 nm. These calibration lamp uncertainties include the calibration uncertainty and the variation resulting from the repeated measurements made before and after all scandium scans. The uncertainties of the radiative lifetimes are given in Table 2. Finally, we were not able to observe the weakest lines from the investigated level. However, we included their contributions as residuals with derived theoretical  $BF$ s from our calculations. The residual  $BF$ s are less than 7 per cent for all levels. The uncertainties in the residuals were estimated to 50 per cent and included in the error budget. The final uncertainties in the oscillator strengths are presented in Table 4 and were derived from the individual contributions using the method described by Sikström et al. (2002).

#### 4 RADIATIVE PARAMETER CALCULATIONS

To calculate branching fractions and the oscillator strengths in Sc II, we used the relativistic Hartree–Fock (HFR) method implemented in the Cowan’s suite of atomic structure computer codes (Cowan 1981). It is modified by including a pseudo-potential and a correction to the electric dipole operator which take into account the core-polarization effects giving rise to the HFR+CPOL technique (Quinet et al. 1999).

In this study, the valence–valence correlation was included using the following configuration interaction (CI) expansions: 3d4s + 3d5s + 3d6s + 3d7s + 3d<sup>2</sup> + 3d4d + 3d5d + 3d6d + 3d7d + 3d5g + 3d6g + 3d7g + 4s<sup>2</sup> + 4s5s + 4s6s + 4s7s + 4s4d + 4s5d + 4s6d + 4s7d + 4s5g + 4s6g + 4s7g + 4p<sup>2</sup> + 4d<sup>2</sup> + 4f<sup>2</sup> + 4p4f for the even parity and 3d4p + 3d5p + 3d6p + 3d7p + 3d4f + 3d5f + 3d6f + 3d7f + 3d6h + 3d7h + 4s4p + 4s5p + 4s6p + 4s7p + 4s4f + 4s5f + 4s6f + 4s7f + 4s6h + 4s7h + 4p4d + 4d4f for the odd parity.

Regarding the core-polarization effects, a Sc IV 3p<sup>6</sup> closed-subshell ionic core was considered where the dipole polarizability,  $\alpha_d = 2.129 a_0^3$ , was taken from the relativistic random-phase approximation calculations of Johnson, Kolb & Huang (1983) and a cut-off radius of 1.17  $a_0$  was estimated as the HFR mean radius of the outermost 3p orbital,  $\langle 3p|r|3p \rangle_{\text{HFR}}$ .

During a least-squares fit procedure, we adjusted some radial integrals to minimize the discrepancies between the Hamiltonian eigenvalues and the experimental energy levels taken from the NIST Atomic Spectra Database (Kramida et al. 2015). The latter are based on the term analysis originally carried out by Russell & Meggers (1927) and later revised by Neufeld (1970) and by Johansson & Litzén (1980). There are 168 levels belonging to the configurations 3d4s, 3d<sup>2</sup>, 3d4p, 4s4p, 3d5s, 3d4d, 3d5p, 4p<sup>2</sup>, 3d4f, 3d6s, 4s5s, 3d5d, 4s4d, 3d5f, 3d5g, 3d7s, 3d6d and 3d6f. The average energies,  $E_{\text{av}}$ , of the above-mentioned known configurations along with their direct,  $F^k$ , exchange,  $G^k$ , electrostatic and spin-orbit,  $\zeta$ , radial parameters were considered in the fit of the energy levels. The *ab initio* and fitted parameter values are reported in Tables 4 and 5 for the even and odd configurations, respectively. The spin-orbit integrals not presented in these tables were fixed to their HFR+CPOL values. The other Slater integrals, including the CI  $R^k$  parameters, not reported here, were fixed to 80 per cent of their *ab initio* values to account for missing CI effects (Cowan 1981). The average deviations of the least-squares fits were 157 cm<sup>-1</sup> for the 93 even-parity experimental levels and 65 cm<sup>-1</sup> for the 75 odd-parity experimental levels.

#### 5 RESULTS AND DISCUSSION

Table 2 compares our TR-LIF and HFR+CPOL lifetimes with other experimental values from the literature (Buchta et al. 1971; Arnesen et al. 1976; Palenius et al. 1976; Vogel et al. 1985; Marsden et al. 1988), the HFR values calculated by Kurucz (2011) and the lifetimes deduced from the semi-empirical oscillator strengths calculated by Ruczkowski et al. (2014). On average, our HFR+CPOL lifetimes are shorter than the measurements for the odd-parity levels and longer for the even-parity levels. The discrepancies range from a few per cent to about 20 per cent, except for the even-parity levels 4p<sup>2</sup> <sup>1</sup>D<sub>2</sub> and 3d6s <sup>3</sup>D<sub>3</sub> where they reach 57 per cent and 49 per cent, respectively. In the former case, this state is strongly mixed (our calculation gives 36 per cent 4p<sup>2</sup> <sup>1</sup>D<sub>2</sub> + 36 per cent 4s4d <sup>1</sup>D<sub>2</sub> + 23 per cent 3d6s <sup>1</sup>D<sub>2</sub>) and an important decay channel (4p<sup>2</sup> <sup>1</sup>D<sub>2</sub> → 3d4p <sup>1</sup>D<sub>2</sub><sup>o</sup>  $BF = 0.0713$ ) is

**Table 4.** Radial parameters adopted in the HFR+CPOL calculations for the even-parity configurations of Sc II. The parameters not listed here have been fixed to their *ab initio* values or to 80 per cent of their HFR+CPOL values for the electrostatic integrals.

Config.	Parameter	<i>Ab initio</i> (cm <sup>-1</sup> )	Fitted (cm <sup>-1</sup> )	Ratio	Note <sup>a</sup>
3d4s	$E_{av}$	1075	1137		
	$\zeta_{3d}$	83	72	0.87	
	$G^2(3d4s)$	11 351	9883	0.87	
3d5s	$E_{av}$	57 881	58 144		
	$\zeta_{3d}$	87	79	0.91	
	$G^2(3d5s)$	2071	1851	0.89	
3d6s	$E_{av}$	77 397	77 497		
	$\zeta_{3d}$	88	82	0.93	
	$G^2(3d6s)$	789	631	0.80	F
3d7s	$E_{av}$	86 487	86 549		
	$\zeta_{3d}$	88	69	0.78	
	$G^2(3d7s)$	393	314	0.80	F
3d <sup>2</sup>	$E_{av}$	11 721	9531		
	$F^2(3d3d)$	49 657	37 346	0.75	
	$F^4(3d3d)$	30 556	22 011	0.72	
	$\alpha$	0	64		
	$\beta$	0	-962		
	$T$	0	3		
	$\zeta_{3d}$	65	59	0.91	
	$\zeta_{4f}$	0	0	1.00	F
3d4d	$E_{av}$	62 210	62 852		
	$\zeta_{3d}$	87	79	0.91	
	$\zeta_{4d}$	8	8	1.00	F
	$F^2(3d4d)$	7539	5977	0.79	
	$F^4(3d4d)$	3599	2816	0.78	
	$G^0(3d4d)$	6862	2467	0.36	
	$G^2(3d4d)$	4352	3238	0.74	
	$G^4(3d4d)$	2927	2327	0.80	
	$\zeta_{3d}$	87	86	0.99	
	$\zeta_{5d}$	3	3	1.00	F
3d5d	$E_{av}$	79 393	79 170		
	$\zeta_{3d}$	87	86	0.99	
	$\zeta_{5d}$	3	3	1.00	F
	$F^2(3d5d)$	2896	2158	0.75	
	$F^4(3d5d)$	1388	1099	0.79	
	$G^0(3d5d)$	2416	1008	0.42	R
	$G^2(3d5d)$	1640	684	0.42	R
	$G^4(3d5d)$	1122	469	0.42	R
3d6d	$E_{av}$	87 550	87 894		
	$\zeta_{3d}$	88	88	1.00	F
	$\zeta_{6d}$	2	2	1.00	F
	$F^2(3d6d)$	1458	1166	0.80	F
	$F^4(3d6d)$	705	564	0.80	F
	$G^0(3d6d)$	1176	941	0.80	F
	$G^2(3d6d)$	822	658	0.80	F
	$G^4(3d6d)$	571	454	0.80	F
	$\zeta_{3d}$	88	78	0.89	
	$\zeta_{5g}$	0	0	1.00	F
3d5g	$E_{av}$	85 492	85 761		
	$\zeta_{3d}$	88	78	0.89	
	$\zeta_{5g}$	0	0	1.00	F
	$F^2(3d5g)$	465	420	0.90	
	$F^4(3d5g)$	42	34	0.81	
	$G^2(3d5g)$	6	5	0.80	F
	$G^4(3d5g)$	4	3	0.80	F
4s <sup>2</sup>	$E_{av}$	16 845	16 876		
	$G^0(4s5s)$	2341	1765	0.75	
4s4d	$E_{av}$	83 034	82 930		
	$\zeta_{4d}$	9	9	1.00	F
4p <sup>2</sup>	$E_{av}$	77 789	80 625		
	$F^2(4p4p)$	28 516	29 802	1.05	
	$\zeta_{4p}$	199	253	1.27	

Note. <sup>a</sup>F and R stand for, respectively, a fixed parameter value and a fixed ratio between these parameters.

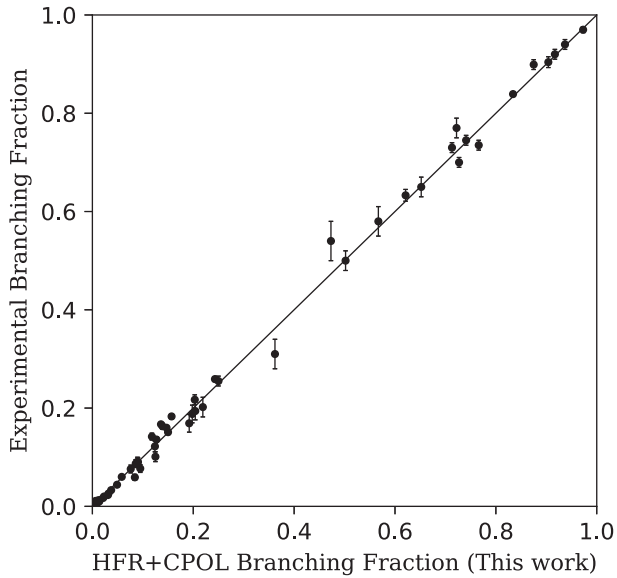
**Table 5.** Radial parameters adopted in the HFR+CPOL calculations for the odd-parity configurations of Sc II. The parameters not listed here have been fixed to their *ab initio* values or to 80 per cent of their HFR+CPOL values for the electrostatic integrals.

Config.	Parameter	<i>Ab initio</i> (cm <sup>-1</sup> )	Fitted (cm <sup>-1</sup> )	Ratio	Note <sup>a</sup>
3d4p	$E_{av}$	28 207	28 996		
	$\zeta_{3d}$	85	91	1.07	
	$\zeta_{4p}$	146	162	1.11	
	$F^2(3d4p)$	14 647	12 024	0.82	
	$G^1(3d4p)$	6709	6289	0.94	
3d5p	$E_{av}$	66 759	66 915		
	$\zeta_{3d}$	87	73	0.84	
	$\zeta_{5p}$	50	50	1.00	F
	$F^2(3d5p)$	4168	3375	0.81	
	$G^1(3d5p)$	1560	1397	0.90	
3d4f	$E_{av}$	75 021	75 609		
	$\zeta_{3d}$	88	74	0.84	
	$\zeta_{4f}$	0	0	1.00	F
	$F^2(3d4f)$	2127	1766	0.83	
	$F^4(3d4f)$	514	367	0.71	
3d5f	$E_{av}$	85 220	85 564		
	$\zeta_{3d}$	88	91	1.03	
	$\zeta_{5f}$	0	0	1.00	F
	$F^2(3d5f)$	1051	841	0.80	F
	$F^4(3d5f)$	296	238	0.80	F
3d6f	$E_{av}$	90 728	91 031		
	$\zeta_{3d}$	88	88	1.00	F
	$\zeta_{6f}$	0	0	1.00	F
	$F^2(3d6f)$	597	478	0.80	F
	$F^4(3d6f)$	181	145	0.80	F
4s4p	$E_{av}$	41 287	43 719		
	$\zeta_{4p}$	197	242	1.23	
	$G^1(4s4p)$	37 326	27 686	0.74	
	$G^3(3d5f)$	168	135	0.80	F
	$G^5(3d5f)$	116	93	0.80	F
$G^3(3d6f)$	111	88	0.80	F	
$G^5(3d6f)$	76	61	0.80	F	

Note. <sup>a</sup>F stands for a fixed parameter value.

affected by cancellation (the cancellation factor as defined by Cowan (1981) is less than 5 per cent) that could explain the overestimated lifetime. Concerning 3d6s <sup>3</sup>D<sub>3</sub> level, no such argument could explain the observed disagreement. The beam-foil measurements of Buchta et al. (1971) can be rejected for the levels 3d4p <sup>3</sup>D<sub>3</sub><sup>o</sup>, <sup>1</sup>P<sub>1</sub><sup>o</sup><sup>1</sup>F<sub>3</sub> as previously stated by Marsden et al. (1988) due to blending problems.

The calculations by Kurucz (2011) show roughly the same systematic discrepancy with experiment (lifetimes shorter for the odd parity and longer for the even parity) as our HFR+CPOL calculations. Although the calculation of Kurucz (2011) shows a better agreement than HFR+CPOL for certain 3d4d levels (<sup>3</sup>F<sub>2,4</sub>, <sup>1</sup>D<sub>2</sub> and <sup>3</sup>P<sub>2</sub>), it does not solve the theory-experiment disagreements observed for the levels 4p<sup>2</sup> <sup>1</sup>D<sub>2</sub> and 3d6s <sup>3</sup>D<sub>3</sub>. The parametric calculation of Ruczkowski et al. (2014) agrees with our HFR+CPOL model within 10 per cent including all levels. Unfortunately, no lifetime value can be deduced from Ruczkowski et al. (2014) for the levels 4p<sup>2</sup> <sup>1</sup>D<sub>2</sub> and 3d6s <sup>3</sup>D<sub>3</sub>. Concerning the level 3d4d <sup>3</sup>G<sub>3</sub>, our



**Figure 2.** A comparison between the HFR+CPOL branching fractions of this work and the experimental values of Lawler & Dakin (1989). The straight line of equality has been drawn.

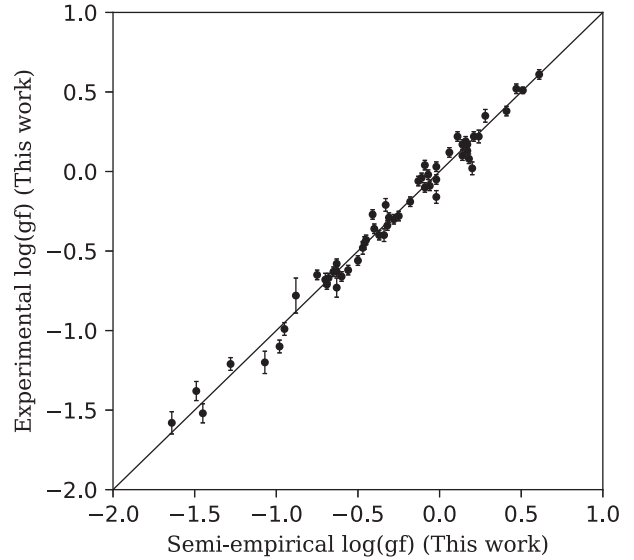
TR-LIF measurement is slightly lower than the one of Marsden et al. (1988) although the error bars do overlap.

For all 3d4p levels, our HFR+CPOL model and the parametric calculation of Ruczkowski et al. (2014) are closer to the measurement of Marsden et al. (1988). The excellent agreement between Marsden et al. (1988) and Ruczkowski et al. (2014) is not surprising as the latter adjusted the dipole transition integrals to the oscillator strengths determined from the branching fraction measurements of Lawler & Dakin (1989) combined with the lifetime measurements of Marsden et al. (1988). For most of the higher levels, the lifetimes calculated by Kurucz (2011) are closer to our measurements than those of Ruczkowski et al. (2014).

Although there is a systematic discrepancy between the theoretical and experimental lifetimes, we find a better agreement when comparing our calculated  $BF$ s with the experimental values. For the high excitation lines, measured in this work, the averaged  $BF$  ratio is  $1.02 \pm 0.16$  with respect to the calculated values. Similarly, Fig. 2 shows the good agreement between  $BF$ s computed in this study using the HFR+CPOL method and the measurements by Lawler & Dakin (1989). Here, the averaged  $BF$  ratio is  $0.98 \pm 0.20$ . Based on these comparisons, the calculated  $BF$ s were combined with our TR-LIF lifetimes and those of Marsden et al. (1988) to determine rescaled transition probabilities and oscillator strengths.

In Table 3, we present our experimental  $\log(gf)$  values, together with the measured  $BF$ s, the uncertainties and the corresponding rescaled theoretical oscillator strengths,  $\log(gf)_{\text{resc}}$ . Fig. 3 illustrates the final agreement between our experimental  $\log(gf)$  values and the calculated  $\log(gf)_{\text{resc}}$ . Table 6 summarizes our calculated radiative parameters along with the weighted transition probabilities ( $gA$ ), the weighted oscillator strengths in the log scale ( $\log(gf)$ ), the HFR+CPOL branching fractions ( $BF$ ), and the cancellation factor ( $CF$ ) as defined by Cowan (1981).

Our rescaled theoretical oscillator strengths are compared to the semi-empirical values calculated by Ruczkowski et al. (2014) in Fig. 4. As expected, the scatter increases for the weak lines, i.e. the transitions with  $\log(gf) \lesssim -1$ , where cancellation effects could be an issue. For instance, the transition  $3d4p\ ^3P_2 - 4p^2\ ^3P_2$  labelled in



**Figure 3.** A comparison between the oscillator strengths determined by the combination of the HFR+CPOL branching fractions and the TR-LIF lifetimes of this work and the experimental oscillator strengths derived in this work. The straight line of equality has been drawn.

Table 6 76589(e)2 – 29824(o)2 has a very low cancellation factor ( $CF = 0.001$ ) that indicates a strong cancellation effect in our HFR+CPOL line strength calculation. Indeed, the rescaled oscillator strength for that transition is  $\log(gf)_{\text{resc}} = -2.83$  which is three orders of magnitude lower (in the linear scale) than the value predicted by Ruczkowski et al. (2014) ( $\log(gf) = -0.02$ ). On the other hand, a transition for which the cancellation effect in our model is not an issue ( $CF > 0.05$ ) such as  $3d4p\ ^3F_3 - 3d4d\ ^3F_4$  (63529(e)4 – 27602(o)3) has an oscillator strength predicted by Ruczkowski et al. (2014) ( $\log(gf) = -3.35$ ) that is two orders of magnitude lower than our rescaled value ( $\log(gf)_{\text{resc}} = -1.44$ ). This could indicate a strong cancellation effect in their calculation. Unfortunately, they did not estimate any cancellation factors. For the strongest transitions, i.e.  $\log(gf) \gtrsim -1$ , the mean scatter drops to about 20 per cent in the linear scale.

In Fig. 5, our semi-empirical values are compared to the calculation of Kurucz (2011) where a similar global correlation is observed. The mean scatter in this case is also found to be  $\sim 20$  per cent for transitions with  $\log(gf) \gtrsim -1$  and increases for weaker lines. Here again, the cancellation factors are not available in Kurucz's data base (Kurucz 2011). But, for example, our predicted strong line  $3d5p\ ^3F_3 - 3d6s\ ^3D_3$  (77387(e)3 – 66564(o)3) with  $\log(gf)_{\text{resc}} = 0.16$  and  $CF = 0.379$  is certainly affected by a strong cancellation effect in the calculation of Kurucz (2011) dramatically lowering its oscillator strength to  $\log(gf) = -2.56$ .

Based on the differences between different sets of  $BF$ s discussed above and including the uncertainties of the experimental lifetimes, we estimate the accuracy of the rescaled theoretical  $f$ -values to be 10 per cent for the strong lines and 15–20 per cent for other lines.

## 6 CONSEQUENCE ON THE SOLAR ABUNDANCE OF SCANDIUM

Scott et al. (2015) have redetermined the solar abundances of the iron-peak elements employing a 3D model atmosphere that takes into account departures from the local thermodynamic equilibrium. However, the significant discrepancy between the photospheric and



**Table 6.** Calculated branching fractions ( $BF$ s), oscillator strengths ( $\log(gf)$ ) and transition probabilities ( $gA$ ) along with the corresponding scaled values ( $\log(gf)_{\text{resc}}$ ,  $gA_{\text{resc}}$ ) in Sc II. Only the transitions depopulating the levels for which the lifetime has been measured are listed. The experimental lifetimes ( $\tau_u$ ) used to scale the  $f$ - and  $A$ -values are also reported.

Upper level <sup>a</sup>		$\tau_u$ (ns)	Lower level <sup>a</sup>	$\lambda^b$ (nm)	$BF$	$gA$ (s <sup>-1</sup> )	$gA_{\text{resc}}$ (s <sup>-1</sup> )	$\log(gf)$	$\log(gf)_{\text{resc}}$	CF <sup>c</sup>						
26081	(o)	2	7.5 <sup>d</sup>	0	(e) 1	383.307	5.91E-03	4.44E+06	3.94E+06	-2.01	-2.06	0.468				
				68	(e) 2	384.305	1.40E-02	1.05E+07	9.32E+06	-1.64	-1.69	0.927				
				178	(e) 3	385.938	1.65E-05	1.24E+04	1.10E+04	-4.56	-4.61	0.006				
				2541	(e) 2	424.682	9.73E-01	7.31E+08	6.49E+08	0.29	0.24	0.975				
				4803	(e) 2	469.827	7.55E-04	5.67E+05	5.03E+05	-2.73	-2.78	0.260				
				4884	(e) 3	471.616	3.77E-05	2.83E+04	2.51E+04	-4.03	-4.08	0.047				
				10945	(e) 2	660.460	6.19E-03	4.65E+06	4.13E+06	-1.53	-1.57	0.036				
				12 102	(e) 1	715.119	6.88E-07	5.17E+02	4.59E+02	-5.41	-5.45	0.009				
				12 154	(e) 2	717.836	8.89E-06	6.68E+03	5.93E+03	-4.30	-4.34	0.008				
				27444	(o)	2	6.2 <sup>d</sup>	0	(e) 1	364.278	7.27E-01	6.40E+08	5.87E+08	0.11	0.07	0.888
								68	(e) 2	365.180	1.57E-01	1.38E+08	1.26E+08	-0.56	-0.60	0.894
178	(e) 3	366.653	6.61E-03					5.82E+06	5.33E+06	-1.93	-1.97	0.927				
2541	(e) 2	401.448	1.27E-02					1.12E+07	1.03E+07	-1.57	-1.61	0.837				
4803	(e) 2	441.556	8.94E-02					7.87E+07	7.21E+07	-0.64	-0.68	0.971				
4884	(e) 3	443.135	7.01E-03					6.17E+06	5.65E+06	-1.74	-1.78	0.466				
10945	(e) 2	605.924	1.13E-04					9.98E+04	9.15E+04	-3.26	-3.30	0.046				
12 102	(e) 1	651.617	2.08E-05					1.83E+04	1.68E+04	-3.93	-3.97	0.676				
12 154	(e) 2	653.872	2.60E-06					2.29E+03	2.10E+03	-4.83	-4.87	0.053				
27602	(o)	3	6.1 <sup>d</sup>					68	(e) 2	363.074	7.66E-01	9.54E+08	8.79E+08	0.28	0.24	0.871
								178	(e) 3	364.531	1.36E-01	1.69E+08	1.56E+08	-0.47	-0.51	0.952
				2541	(e) 2	398.906	7.51E-04	9.35E+05	8.62E+05	-2.65	-2.69	0.768				
				4803	(e) 2	438.481	7.95E-03	9.90E+06	9.12E+06	-1.54	-1.58	0.966				
				4884	(e) 3	440.039	8.59E-02	1.07E+08	9.86E+07	-0.51	-0.54	0.975				
				4988	(e) 4	442.067	3.37E-03	4.20E+06	3.87E+06	-1.91	-1.95	0.242				
				10945	(e) 2	600.150	4.67E-06	5.82E+03	5.36E+03	-4.50	-4.54	0.520				
				12 154	(e) 2	647.153	4.47E-05	5.57E+04	5.13E+04	-3.46	-3.49	0.612				
				14 261	(e) 4	749.355	5.08E-08	6.33E+01	5.83E+01	-6.28	-6.31	0.089				
				27841	(o)	4	6.1 <sup>d</sup>	178	(e) 3	361.383	9.04E-01	1.47E+09	1.33E+09	0.46	0.42	0.949
								4884	(e) 3	435.460	6.12E-03	9.96E+06	9.03E+06	-1.55	-1.59	0.976
4988	(e) 4	437.446	9.04E-02					1.47E+08	1.33E+08	-0.37	-0.42	0.976				
14 261	(e) 4	736.173	7.87E-07					1.28E+03	1.16E+03	-4.99	-5.03	0.912				
27918	(o)	1	4.7 <sup>d</sup>	0	(e) 1	358.092	5.67E-01	3.83E+08	3.62E+08	-0.13	-0.16	0.931				
				68	(e) 2	358.963	2.03E-01	1.37E+08	1.29E+08	-0.58	-0.60	0.950				
				2541	(e) 2	393.949	6.08E-06	4.11E+03	3.88E+03	-5.02	-5.04	0.001				
				4803	(e) 2	432.500	2.19E-01	1.48E+08	1.40E+08	-0.38	-0.41	0.961				
				10945	(e) 2	589.000	1.73E-04	1.17E+05	1.11E+05	-3.21	-3.24	0.214				
				11 736	(e) 0	617.822	4.66E-04	3.15E+05	2.98E+05	-2.74	-2.77	0.704				
				12 074	(e) 0	630.992	6.22E-03	4.20E+06	3.97E+06	-1.60	-1.63	0.611				
				12 102	(e) 1	632.085	4.22E-03	2.85E+06	2.69E+06	-1.77	-1.79	0.563				
				12 154	(e) 2	634.207	2.25E-04	1.52E+05	1.44E+05	-3.04	-3.06	0.323				
				25 955	(e) 0	5093.945	6.48E-08	4.38E+01	4.14E+01	-4.72	-4.79	0.157				
				28021	(o)	2	4.7 <sup>d</sup>	0	(e) 1	356.770	1.39E-01	1.57E+08	1.48E+08	-0.52	-0.55	0.950
68	(e) 2	357.634	5.02E-01					5.68E+08	5.34E+08	0.04	0.01	0.857				
178	(e) 3	359.047	1.27E-01					1.44E+08	1.35E+08	-0.55	-0.58	0.923				
2541	(e) 2	392.348	2.16E-03					2.44E+06	2.29E+06	-2.25	-2.28	0.741				
4803	(e) 2	430.571	2.13E-02					2.41E+07	2.26E+07	-1.17	-1.20	0.721				
4884	(e) 3	432.073	1.98E-01					2.24E+08	2.10E+08	-0.20	-0.23	0.963				
10945	(e) 2	585.430	8.92E-07					1.01E+03	9.49E+02	-5.28	-5.31	0.002				
12 102	(e) 1	627.975	8.83E-03					1.00E+07	9.40E+06	-1.23	-1.25	0.651				
12 154	(e) 2	630.070	2.24E-03					2.54E+06	2.39E+06	-1.82	-1.85	0.443				
28161	(o)	3	4.7 <sup>d</sup>					68	(e) 2	355.853	1.18E-01	1.88E+08	1.76E+08	-0.45	-0.48	0.968
								178	(e) 3	357.253	6.52E-01	1.04E+09	9.72E+08	0.30	0.27	0.894
				2541	(e) 2	390.206	2.31E-05	3.69E+04	3.45E+04	-4.08	-4.10	0.023				
				4803	(e) 2	427.993	3.06E-04	4.88E+05	4.56E+05	-2.87	-2.90	0.383				
				4884	(e) 3	429.477	1.37E-02	2.18E+07	2.04E+07	-1.22	-1.25	0.601				
				4988	(e) 4	431.408	2.04E-01	3.26E+08	3.05E+08	-0.04	-0.07	0.963				
				10945	(e) 2	580.673	1.57E-04	2.50E+05	2.34E+05	-2.90	-2.93	0.555				
				12 154	(e) 2	624.564	1.10E-02	1.76E+07	1.64E+07	-0.99	-1.02	0.638				
				14 261	(e) 4	719.234	3.06E-05	4.88E+04	4.56E+04	-3.42	-3.45	0.419				
				29736	(o)	0	7.7 <sup>d</sup>	0	(e) 1	336.193	8.75E-01	1.38E+08	1.14E+08	-0.63	-0.72	0.519
								12 102	(e) 1	566.904	1.25E-01	1.98E+07	1.63E+07	-1.02	-1.10	0.885

Table 6 – continued

Upper level <sup>d</sup>	$\tau_u$ (ns)	Lower level <sup>a</sup>	$\lambda^b$ (nm)	$BF$	$gA$ (s <sup>-1</sup> )	$gA_{\text{resc}}$ (s <sup>-1</sup> )	$\log(gf)$	$\log(gf)_{\text{resc}}$	$CF^c$						
29 742	(o)	1	7.6 <sup>d</sup>	0 (e) 1	336.127	2.43E-01	1.14E+08	9.57E+07	-0.72	-0.79	0.538				
				68 (e) 2	336.894	6.21E-01	2.92E+08	2.45E+08	-0.30	-0.38	0.485				
				2541 (e) 2	367.526	4.00E-03	1.88E+06	1.58E+06	-2.42	-2.49	0.070				
				4803 (e) 2	400.860	2.21E-04	1.04E+05	8.73E+04	-3.60	-3.68	0.479				
				10 945 (e) 2	531.835	8.60E-03	4.04E+06	3.39E+06	-1.77	-1.84	0.723				
				11 736 (e) 0	555.222	5.30E-03	2.49E+06	2.09E+06	-1.94	-2.01	0.771				
				12 074 (e) 0	565.836	3.72E-02	1.75E+07	1.47E+07	-1.08	-1.15	0.736				
				12 102 (e) 1	566.715	3.19E-02	1.50E+07	1.26E+07	-1.15	-1.22	0.881				
				12 154 (e) 2	568.420	4.89E-02	2.30E+07	1.93E+07	-0.96	-1.03	0.813				
				25 955 (e) 0	2639.920	7.02E-06	3.30E+03	2.77E+03	-3.46	-3.54	0.175				
				29 824	(o)	2	7.4 <sup>d</sup>	0 (e) 1	335.205	1.05E-02	8.36E+06	7.12E+06	-1.85	-1.92	0.529
								68 (e) 2	335.968	1.47E-01	1.17E+08	9.96E+07	-0.71	-0.77	0.537
								178 (e) 3	337.215	7.13E-01	5.66E+08	4.82E+08	-0.02	-0.09	0.506
2541 (e) 2	366.425	2.42E-03	1.92E+06					1.63E+06	-2.42	-2.48	0.731				
4803 (e) 2	399.550	3.24E-05	2.57E+04					2.19E+04	-4.21	-4.28	0.865				
4884 (e) 3	400.843	1.69E-04	1.34E+05					1.14E+05	-3.49	-3.56	0.912				
10 945 (e) 2	529.531	4.02E-04	3.19E+05					2.72E+05	-2.88	-2.94	0.210				
12 102 (e) 1	564.100	3.08E-02	2.44E+07					2.08E+07	-0.94	-1.00	0.825				
12 154 (e) 2	565.790	9.49E-02	7.53E+07					6.41E+07	-0.45	-0.51	0.879				
30 816	(o)	1	8.8 <sup>d</sup>					0 (e) 1	324.416	7.98E-04	2.95E+05	2.72E+05	-3.33	-3.37	0.025
				68 (e) 2	325.131	2.31E-02	8.56E+06	7.89E+06	-1.87	-1.90	0.259				
				2541 (e) 2	353.571	4.73E-01	1.75E+08	1.61E+08	-0.49	-0.52	0.193				
				4803 (e) 2	384.317	2.68E-04	9.91E+04	9.14E+04	-3.66	-3.69	0.065				
				10 945 (e) 2	503.102	3.62E-01	1.34E+08	1.24E+08	-0.29	-0.33	0.723				
				11 736 (e) 0	523.981	1.24E-01	4.60E+07	4.24E+07	-0.72	-0.76	0.685				
				12 074 (e) 0	533.424	8.14E-03	3.01E+06	2.77E+06	-1.89	-1.93	0.773				
				12 102 (e) 1	534.205	7.73E-04	2.86E+05	2.64E+05	-2.91	-2.95	0.422				
				12 154 (e) 2	535.720	6.06E-03	2.24E+06	2.06E+06	-2.01	-2.05	0.758				
				25 955 (e) 0	2056.840	8.71E-04	3.22E+05	2.97E+05	-1.67	-1.73	0.214				
				32 350	(o)	3	5.1 <sup>d</sup>	68 (e) 2	309.678	5.40E-04	8.08E+05	7.41E+05	-2.94	-2.97	0.182
178 (e) 3	310.737	5.99E-04	8.97E+05					8.22E+05	-2.89	-2.92	0.835				
2541 (e) 2	335.372	7.22E-01	1.08E+09					9.90E+08	0.26	0.22	0.640				
4803 (e) 2	362.911	2.65E-04	3.96E+05					3.63E+05	-3.11	-3.14	0.722				
4884 (e) 3	363.977	1.13E-05	1.69E+04					1.55E+04	-4.47	-4.51	0.411				
4988 (e) 4	365.364	9.09E-05	1.36E+05					1.25E+05	-3.57	-3.60	0.222				
10 945 (e) 2	467.041	8.42E-02	1.26E+08					1.16E+08	-0.39	-0.42	0.625				
12 154 (e) 2	495.020	4.18E-04	6.25E+05					5.73E+05	-2.64	-2.68	0.621				
14 261 (e) 4	552.679	1.92E-01	2.88E+08					2.64E+08	0.12	0.08	0.917				
39 002	(o)	0	3.7 <sup>d</sup>					0 (e) 1	256.319	9.94E-01	2.70E+08	2.69E+08	-0.58	-0.58	0.958
				12 102 (e) 1	371.632	5.85E-03	1.59E+06	1.58E+06	-2.48	-2.48	0.139				
39 115	(o)	1	3.7 <sup>d</sup>	0 (e) 1	255.580	2.50E-01	2.04E+08	2.03E+08	-0.70	-0.70	0.958				
				68 (e) 2	256.023	7.41E-01	6.04E+08	6.01E+08	-0.23	-0.23	0.955				
				2541 (e) 2	273.337	4.16E-04	3.39E+05	3.37E+05	-3.42	-3.42	0.244				
				4803 (e) 2	291.357	4.23E-07	3.45E+02	3.43E+02	-6.36	-6.36	0.362				
				10 945 (e) 2	354.880	2.80E-05	2.28E+04	2.27E+04	-4.37	-4.37	0.095				
				11 736 (e) 0	365.144	5.13E-07	4.18E+02	4.16E+02	-6.08	-6.08	0.001				
				12 074 (e) 0	369.704	1.94E-03	1.58E+06	1.57E+06	-2.49	-2.49	0.137				
				12 102 (e) 1	370.079	1.44E-03	1.17E+06	1.16E+06	-2.62	-2.62	0.137				
39 346	(o)	2	3.8 <sup>d</sup>	0 (e) 1	254.082	1.02E-02	1.38E+07	1.34E+07	-1.88	-1.89	0.957				
				68 (e) 2	254.520	1.50E-01	2.04E+08	1.98E+08	-0.70	-0.72	0.957				
				178 (e) 3	255.235	8.34E-01	1.13E+09	1.10E+09	0.04	0.03	0.956				
				2541 (e) 2	271.625	2.29E-04	3.10E+05	3.01E+05	-3.47	-3.48	0.829				
				4803 (e) 2	289.412	1.49E-07	2.02E+02	1.96E+02	-6.60	-6.61	0.745				
				4884 (e) 3	290.090	4.46E-07	6.05E+02	5.87E+02	-6.12	-6.13	0.917				
				10 945 (e) 2	352.000	2.32E-05	3.14E+04	3.05E+04	-4.23	-4.25	0.097				
				12 102 (e) 1	366.949	1.36E-03	1.85E+06	1.80E+06	-2.43	-2.44	0.129				
				12 154 (e) 2	367.663	4.12E-03	5.58E+06	5.42E+06	-1.95	-1.96	0.132				
				57 744	(e)	3	3.20 <sup>e</sup>	26 081 (o) 2	315.739	5.95E-06	1.19E+04	1.30E+04	-4.75	-4.71	0.004
27 444 (o) 2	329.936	1.75E-03	3.49E+06					3.82E+06	-2.25	-2.21	0.660				
27 602 (o) 3	331.673	4.87E-02	9.74E+07					1.07E+08	-0.80	-0.75	0.694				
27 841 (o) 4	334.323	3.77E-01	7.53E+08					8.24E+08	0.10	0.14	0.668				

Table 6 – continued

Upper level <sup>a</sup>	$\tau_u$ (ns)	Lower level <sup>a</sup>	$\lambda^b$ (nm)	$BF$	$gA$ (s <sup>-1</sup> )	$gA_{\text{resc}}$ (s <sup>-1</sup> )	$\log(gf)$	$\log(gf)_{\text{resc}}$	CF <sup>c</sup>								
58 252	(e)	2	3.26 <sup>e</sup>	28 021	(o)	2	336.347	5.55E-02	1.11E+08	1.21E+08	-0.73	-0.69	0.851				
				28 161	(o)	3	337.938	3.93E-01	7.86E+08	8.60E+08	0.13	0.17	0.834				
				29 824	(o)	2	358.064	1.24E-01	2.47E+08	2.70E+08	-0.32	-0.28	0.458				
				32 350	(o)	3	393.683	8.95E-05	1.79E+05	1.96E+05	-3.38	-3.34	0.357				
				39 346	(o)	2	543.375	5.65E-04	1.13E+06	1.24E+06	-2.30	-2.26	0.027				
				26 081	(o)	2	310.751	3.70E-01	4.99E+08	5.67E+08	-0.14	-0.09	0.549				
				27 444	(o)	2	324.493	9.48E-03	1.28E+07	1.45E+07	-1.70	-1.64	0.583				
				27 602	(o)	3	326.174	7.63E-03	1.03E+07	1.17E+07	-1.79	-1.73	0.542				
				27 918	(o)	1	329.565	3.92E-04	5.29E+05	6.01E+05	-3.07	-3.01	0.062				
				28 021	(o)	2	330.693	1.11E-02	1.50E+07	1.70E+07	-1.61	-1.55	0.754				
				28 161	(o)	3	332.231	7.48E-04	1.01E+06	1.15E+06	-2.78	-2.72	0.136				
				29 742	(o)	1	350.655	4.07E-04	5.50E+05	6.25E+05	-2.99	-2.94	0.016				
				29 824	(o)	2	351.663	1.91E-03	2.58E+06	2.93E+06	-2.32	-2.26	0.439				
				30 816	(o)	1	364.376	1.56E-01	2.10E+08	2.39E+08	-0.38	-0.32	0.524				
				32 350	(o)	3	385.960	4.43E-01	5.98E+08	6.79E+08	0.13	0.18	0.811				
				39 115	(o)	1	522.401	6.04E-07	8.16E+02	9.27E+02	-5.48	-5.42	0.001				
59 528	(e)	3	2.32 <sup>e</sup>	39 346	(o)	2	528.770	5.48E-06	7.40E+03	8.41E+03	-4.51	-4.45	0.032				
				55 715	(o)	1	3941.008	1.07E-04	1.45E+05	1.65E+05	-1.48	-1.42	0.272				
				26 081	(o)	2	298.893	7.30E-01	1.90E+09	2.20E+09	0.41	0.47	0.813				
				27 444	(o)	2	311.585	5.11E-04	1.33E+06	1.54E+06	-2.71	-2.65	0.017				
				27 602	(o)	3	313.134	1.28E-03	3.34E+06	3.87E+06	-2.31	-2.24	0.483				
				27 841	(o)	4	315.495	1.30E-03	3.38E+06	3.92E+06	-2.30	-2.23	0.444				
				28 021	(o)	2	317.297	8.37E-05	2.18E+05	2.53E+05	-3.48	-3.42	0.013				
				28 161	(o)	3	318.712	7.76E-03	2.02E+07	2.34E+07	-1.51	-1.45	0.493				
				29 824	(o)	2	336.553	2.29E-03	5.97E+06	6.92E+06	-1.99	-1.93	0.229				
				32 350	(o)	3	367.834	2.57E-01	6.69E+08	7.75E+08	0.14	0.20	0.867				
				39 346	(o)	2	495.331	1.85E-04	4.81E+05	5.57E+05	-2.75	-2.69	0.621				
				26 081	(o)	2	295.826	1.55E-02	1.71E+07	2.09E+07	-1.64	-1.56	0.658				
				27 444	(o)	2	308.254	1.22E-01	1.34E+08	1.64E+08	-0.72	-0.63	0.631				
				27 918	(o)	1	312.827	3.93E-01	4.33E+08	5.29E+08	-0.20	-0.11	0.597				
				28 021	(o)	2	313.843	1.12E-01	1.23E+08	1.50E+08	-0.74	-0.65	0.481				
				59 875	(e)	1	2.23 <sup>e</sup>	29 736	(o)	0	331.703	1.92E-01	2.11E+08	2.58E+08	-0.46	-0.37	0.899
29 742	(o)	1	331.768					1.42E-01	1.57E+08	1.92E+08	-0.59	-0.50	0.750				
29 824	(o)	2	332.670					7.03E-03	7.74E+06	9.45E+06	-1.89	-1.80	0.463				
30 816	(o)	1	344.024					9.17E-04	1.01E+06	1.23E+06	-2.75	-2.66	0.021				
39 002	(o)	0	478.957					9.08E-03	1.00E+07	1.22E+07	-1.46	-1.38	0.769				
39 115	(o)	1	481.560					6.81E-03	7.50E+06	9.16E+06	-1.58	-1.50	0.766				
39 346	(o)	2	486.967					3.60E-04	3.97E+05	4.85E+05	-2.85	-2.76	0.542				
55 715	(o)	1	2403.352					2.00E-05	2.20E+04	2.69E+04	-2.72	-2.63	0.500				
26 081	(o)	2	295.351					5.27E-05	9.68E+04	1.14E+05	-3.89	-3.83	0.011				
27 444	(o)	2	307.738					9.53E-03	1.75E+07	2.05E+07	-1.61	-1.53	0.294				
27 602	(o)	3	309.249					1.16E-01	2.13E+08	2.50E+08	-0.52	-0.45	0.667				
27 918	(o)	1	312.296					7.46E-02	1.37E+08	1.61E+08	-0.70	-0.63	0.512				
28 021	(o)	2	313.309					3.62E-01	6.64E+08	7.79E+08	-0.01	0.06	0.578				
28 161	(o)	3	314.688					6.48E-02	1.19E+08	1.40E+08	-0.76	-0.68	0.420				
29 742	(o)	1	331.170					2.71E-01	4.97E+08	5.83E+08	-0.09	-0.02	0.907				
29 824	(o)	2	332.069					7.73E-02	1.42E+08	1.67E+08	-0.63	-0.56	0.690				
59 929	(e)	2	2.32 <sup>e</sup>	30 816	(o)	1	343.382	3.84E-03	7.06E+06	8.29E+06	-1.91	-1.83	0.362				
				32 350	(o)	3	362.485	4.89E-06	8.98E+03	1.05E+04	-4.75	-4.68	0.022				
				39 115	(o)	1	480.302	1.24E-02	2.28E+07	2.68E+07	-1.10	-1.03	0.783				
				39 346	(o)	2	485.680	3.79E-03	6.96E+06	8.17E+06	-1.61	-1.54	0.716				
				55 715	(o)	1	2372.344	1.32E-09	2.42E+00	2.84E+00	-6.69	-6.62	0.003				
				26 081	(o)	2	294.720	5.91E-03	1.50E+07	1.72E+07	-1.71	-1.65	0.241				
				27 444	(o)	2	307.053	9.57E-04	2.43E+06	2.78E+06	-2.47	-2.41	0.144				
				27 602	(o)	3	308.558	3.10E-03	7.86E+06	9.00E+06	-1.95	-1.89	0.088				
				27 841	(o)	4	310.850	1.09E-01	2.76E+08	3.16E+08	-0.40	-0.34	0.677				
				28 021	(o)	2	312.599	5.52E-02	1.40E+08	1.60E+08	-0.69	-0.63	0.476				
				28 161	(o)	3	313.972	4.41E-01	1.12E+09	1.28E+09	0.22	0.28	0.569				
				29 824	(o)	2	331.272	3.63E-01	9.22E+08	1.06E+09	0.18	0.24	0.916				
				32 350	(o)	3	361.535	5.75E-03	1.46E+07	1.67E+07	-1.54	-1.48	0.807				
				39 346	(o)	2	483.977	1.58E-02	4.00E+07	4.58E+07	-0.85	-0.79	0.794				
				26 081	(o)	2	292.433	2.06E-03	5.78E+06	6.59E+06	-2.13	-2.07	0.062				
				27 444	(o)	2	304.572	9.17E-01	2.57E+09	2.93E+09	0.55	0.61	0.862				
27 602	(o)	3	306.052	7.60E-02	2.13E+08	2.43E+08	-0.52	-0.47	0.784								
60 002	(e)	3	2.41 <sup>e</sup>	26 081	(o)	2	294.720	5.91E-03	1.50E+07	1.72E+07	-1.71	-1.65	0.241				
				27 444	(o)	2	307.053	9.57E-04	2.43E+06	2.78E+06	-2.47	-2.41	0.144				
				27 602	(o)	3	308.558	3.10E-03	7.86E+06	9.00E+06	-1.95	-1.89	0.088				
				27 841	(o)	4	310.850	1.09E-01	2.76E+08	3.16E+08	-0.40	-0.34	0.677				
				28 021	(o)	2	312.599	5.52E-02	1.40E+08	1.60E+08	-0.69	-0.63	0.476				
				28 161	(o)	3	313.972	4.41E-01	1.12E+09	1.28E+09	0.22	0.28	0.569				
				29 824	(o)	2	331.272	3.63E-01	9.22E+08	1.06E+09	0.18	0.24	0.916				
				32 350	(o)	3	361.535	5.75E-03	1.46E+07	1.67E+07	-1.54	-1.48	0.807				
				39 346	(o)	2	483.977	1.58E-02	4.00E+07	4.58E+07	-0.85	-0.79	0.794				
				26 081	(o)	2	292.433	2.06E-03	5.78E+06	6.59E+06	-2.13	-2.07	0.062				
				27 444	(o)	2	304.572	9.17E-01	2.57E+09	2.93E+09	0.55	0.61	0.862				
				27 602	(o)	3	306.052	7.60E-02	2.13E+08	2.43E+08	-0.52	-0.47	0.784				
				60 267	(e)	3	2.19 <sup>e</sup>	26 081	(o)	2	292.433	2.06E-03	5.78E+06	6.59E+06	-2.13	-2.07	0.062
								27 444	(o)	2	304.572	9.17E-01	2.57E+09	2.93E+09	0.55	0.61	0.862
								27 602	(o)	3	306.052	7.60E-02	2.13E+08	2.43E+08	-0.52	-0.47	0.784

Table 6 – continued

Upper level <sup>d</sup>	$\tau_u$ (ns)	Lower level <sup>a</sup>	$\lambda^b$ (nm)	$BF$	$gA$ (s <sup>-1</sup> )	$gA_{\text{resc}}$ (s <sup>-1</sup> )	$\log(gf)$	$\log(gf)_{\text{resc}}$	CF <sup>c</sup>	
		27 841 (o) 4	308.307	7.21E-04	2.02E+06	2.30E+06	-2.54	-2.48	0.321	
		28 021 (o) 2	310.027	1.66E-03	4.65E+06	5.30E+06	-2.17	-2.12	0.469	
		28 161 (o) 3	311.378	9.67E-04	2.71E+06	3.09E+06	-2.40	-2.35	0.618	
		29 824 (o) 2	328.386	2.14E-04	5.99E+05	6.83E+05	-3.01	-2.96	0.858	
		32 350 (o) 3	358.100	1.44E-03	4.03E+06	4.60E+06	-2.11	-2.05	0.870	
		39 346 (o) 2	477.840	6.67E-06	1.87E+04	2.13E+04	-4.19	-4.14	0.781	
60 348	(e) 4	2.4 <sup>d</sup>	27 602 (o) 3	305.292	9.37E-01	3.35E+09	3.52E+09	0.67	0.69	0.863
			27 841 (o) 4	307.536	5.82E-02	2.08E+08	2.18E+08	-0.53	-0.51	0.781
			28 161 (o) 3	310.592	4.25E-03	1.52E+07	1.60E+07	-1.66	-1.64	0.846
			32 350 (o) 3	357.060	7.98E-05	2.85E+05	2.99E+05	-3.26	-3.24	0.295
60 400	(e) 1	2.44 <sup>e</sup>	26 081 (o) 2	291.298	2.97E-01	3.08E+08	3.65E+08	-0.41	-0.33	0.746
			27 444 (o) 2	303.340	1.18E-02	1.23E+07	1.46E+07	-1.77	-1.70	0.731
			27 918 (o) 1	307.767	2.40E-03	2.49E+06	2.95E+06	-2.45	-2.38	0.076
			28 021 (o) 2	308.751	4.66E-03	4.84E+06	5.73E+06	-2.16	-2.09	0.499
			29 736 (o) 0	326.020	1.01E-02	1.05E+07	1.24E+07	-1.78	-1.70	0.891
			29 742 (o) 1	326.082	1.70E-02	1.76E+07	2.08E+07	-1.55	-1.48	0.255
			29 824 (o) 2	326.955	3.85E-03	4.00E+06	4.74E+06	-2.19	-2.12	0.314
			30 816 (o) 1	337.915	6.51E-01	6.76E+08	8.01E+08	0.06	0.14	0.880
			39 002 (o) 0	467.198	5.28E-04	5.48E+05	6.49E+05	-2.75	-2.67	0.735
			39 115 (o) 1	469.675	1.63E-05	1.69E+04	2.00E+04	-4.26	-4.18	0.014
			39 346 (o) 2	474.816	3.58E-04	3.72E+05	4.41E+05	-2.90	-2.83	0.590
			55 715 (o) 1	2133.866	1.43E-03	1.48E+06	1.75E+06	-1.00	-0.92	0.554
60 457	(e) 5	2.5 <sup>d</sup>	27 841 (o) 4	306.511	1.00E+00	4.33E+09	4.40E+09	0.79	0.79	0.865
61 071	(e) 1	2.45 <sup>e</sup>	26 081 (o) 2	285.711	6.87E-04	7.30E+05	8.41E+05	-3.05	-2.99	0.084
			27 444 (o) 2	297.287	1.30E-05	1.38E+04	1.59E+04	-4.74	-4.68	0.061
			27 918 (o) 1	301.538	8.24E-05	8.76E+04	1.01E+05	-3.93	-3.86	0.091
			28 021 (o) 2	302.483	6.95E-05	7.39E+04	8.51E+04	-4.00	-3.93	0.056
			29 736 (o) 0	319.038	1.06E-01	1.13E+08	1.30E+08	-0.76	-0.70	0.841
			29 742 (o) 1	319.098	3.00E-01	3.19E+08	3.68E+08	-0.31	-0.25	0.802
			29 824 (o) 2	319.933	5.06E-01	5.38E+08	6.20E+08	-0.08	-0.02	0.837
			30 816 (o) 1	330.421	2.64E-02	2.81E+07	3.24E+07	-1.34	-1.28	0.895
			39 002 (o) 0	452.993	6.99E-03	7.43E+06	8.56E+06	-1.64	-1.58	0.718
			39 115 (o) 1	455.321	2.08E-02	2.21E+07	2.55E+07	-1.17	-1.10	0.750
			39 346 (o) 2	460.151	3.23E-02	3.43E+07	3.95E+07	-0.96	-0.90	0.742
			55 715 (o) 1	1866.531	1.98E-05	2.10E+04	2.42E+04	-2.96	-2.90	0.528
63 375	(e) 2	2.15 <sup>e</sup>	26 081 (o) 2	268.065	1.38E-02	2.88E+07	3.21E+07	-1.51	-1.46	0.561
			27 444 (o) 2	278.230	3.17E-01	6.61E+08	7.37E+08	-0.12	-0.07	0.673
			27 602 (o) 3	279.464	4.12E-02	8.59E+07	9.58E+07	-1.00	-0.95	0.701
			27 918 (o) 1	281.950	5.37E-01	1.12E+09	1.25E+09	0.12	0.17	0.685
			28 021 (o) 2	282.776	8.92E-02	1.86E+08	2.07E+08	-0.65	-0.60	0.516
			28 161 (o) 3	283.899	1.87E-03	3.91E+06	4.36E+06	-2.33	-2.28	0.263
			29 742 (o) 1	297.245	3.20E-05	6.68E+04	7.45E+04	-4.05	-4.01	0.023
			29 824 (o) 2	297.969	2.25E-05	4.70E+04	5.24E+04	-4.20	-4.16	0.075
			30 816 (o) 1	307.046	2.30E-06	4.79E+03	5.34E+03	-5.17	-5.12	0.000
			32 350 (o) 3	322.231	1.81E-04	3.77E+05	4.20E+05	-3.23	-3.18	0.221
			39 115 (o) 1	412.092	1.83E-06	3.81E+03	4.25E+03	-5.01	-4.97	0.218
			39 346 (o) 2	416.045	3.25E-07	6.78E+02	7.56E+02	-5.76	-5.71	0.068
			55 715 (o) 1	1305.250	5.75E-06	1.20E+04	1.34E+04	-3.51	-3.47	0.207
63 529	(e) 4	2.19 <sup>e</sup>	27 602 (o) 3	278.267	7.54E-03	2.79E+07	3.10E+07	-1.49	-1.44	0.119
			27 841 (o) 4	280.130	3.33E-01	1.23E+09	1.37E+09	0.16	0.21	0.704
			28 161 (o) 3	282.663	6.60E-01	2.44E+09	2.71E+09	0.47	0.51	0.697
			32 350 (o) 3	320.641	3.38E-05	1.25E+05	1.39E+05	-3.71	-3.67	0.014
64 367	(e) 2	2.25 <sup>e</sup>	26 081 (o) 2	261.119	5.63E-01	1.03E+09	1.25E+09	0.03	0.11	0.410
			27 444 (o) 2	270.754	1.32E-02	2.41E+07	2.93E+07	-1.58	-1.49	0.444
			27 602 (o) 3	271.923	5.47E-04	1.00E+06	1.21E+06	-2.96	-2.87	0.510
			27 918 (o) 1	274.276	3.89E-04	7.11E+05	8.64E+05	-3.10	-3.01	0.023
			28 021 (o) 2	275.057	4.30E-04	7.87E+05	9.56E+05	-3.05	-2.96	0.021
			28 161 (o) 3	276.120	8.53E-03	1.56E+07	1.90E+07	-1.75	-1.66	0.258
			29 742 (o) 1	288.728	3.03E-02	5.55E+07	6.74E+07	-1.16	-1.07	0.503
			29 824 (o) 2	289.412	1.85E-02	3.38E+07	4.11E+07	-1.37	-1.29	0.305
			30 816 (o) 1	297.967	3.20E-01	5.85E+08	7.11E+08	-0.11	-0.02	0.567
			32 350 (o) 3	312.247	4.32E-02	7.90E+07	9.60E+07	-0.94	-0.85	0.207
			39 115 (o) 1	395.902	3.40E-05	6.22E+04	7.56E+04	-3.84	-3.75	0.017

Table 6 – continued

Upper level <sup>a</sup>				$\tau_u$ (ns)	Lower level <sup>a</sup>			$\lambda^b$ (nm)	$BF$	$gA$ (s <sup>-1</sup> )	$gA_{\text{resc}}$ (s <sup>-1</sup> )	$\log(gf)$	$\log(gf)_{\text{resc}}$	$CF^c$
64 706	(e)	2	2.51 <sup>e</sup>	39 346	(o)	2	399.549	9.57E-05	1.75E+05	2.13E+05	-3.38	-3.29	0.021	
				55 715	(o)	1	1155.577	1.90E-03	3.48E+06	4.23E+06	-1.16	-1.07	0.207	
				26 081	(o)	2	258.825	2.19E-02	3.43E+07	4.36E+07	-1.46	-1.36	0.199	
				27 444	(o)	2	268.289	8.23E-04	1.29E+06	1.64E+06	-2.86	-2.75	0.238	
				27 602	(o)	3	269.437	6.38E-04	9.99E+05	1.27E+06	-2.97	-2.86	0.208	
				27 918	(o)	1	271.746	8.55E-04	1.34E+06	1.70E+06	-2.83	-2.72	0.055	
				28 021	(o)	2	272.513	2.21E-02	3.46E+07	4.40E+07	-1.42	-1.31	0.154	
				28 161	(o)	3	273.556	1.56E-01	2.44E+08	3.10E+08	-0.56	-0.46	0.252	
				29 742	(o)	1	285.927	1.63E-01	5.26E+08	3.26E+08	-0.50	-0.40	0.424	
				29 824	(o)	2	286.597	5.87E-01	9.20E+08	1.17E+09	0.06	0.16	0.545	
				30 816	(o)	1	294.984	4.05E-02	6.35E+07	8.07E+07	-1.08	-0.98	0.562	
				32 350	(o)	3	308.973	3.45E-03	5.40E+06	6.87E+06	-2.11	-2.01	0.222	
				39 115	(o)	1	390.654	7.53E-04	1.18E+06	1.50E+06	-2.57	-2.46	0.021	
				39 346	(o)	2	394.204	2.45E-03	3.84E+06	4.88E+06	-2.05	-1.94	0.024	
74 433	(e)	2	3.80 <sup>e</sup>	55 715	(o)	1	1111.977	1.19E-04	1.87E+05	2.38E+05	-2.46	-2.36	0.198	
				26 081	(o)	2	206.751	7.13E-02	5.98E+07	9.38E+07	-1.42	-1.22	0.047	
				27 444	(o)	2	212.746	1.00E-03	8.40E+05	1.32E+06	-3.25	-3.05	0.054	
				27 602	(o)	3	213.467	5.58E-05	4.68E+04	7.34E+04	-4.51	-4.30	0.083	
				27 918	(o)	1	214.914	8.51E-07	7.14E+02	1.12E+03	-6.32	-6.11	0.000	
				28 021	(o)	2	215.394	2.11E-04	1.77E+05	2.78E+05	-3.92	-3.71	0.024	
				28 161	(o)	3	216.045	1.07E-02	8.94E+06	1.40E+07	-2.21	-2.01	0.455	
				29 742	(o)	1	223.689	5.46E-04	4.58E+05	7.19E+05	-3.47	-3.27	0.005	
				29 824	(o)	2	224.099	2.39E-05	2.00E+04	3.14E+04	-4.83	-4.63	0.001	
				30 816	(o)	1	229.195	4.05E-03	3.40E+06	5.34E+06	-2.58	-2.38	0.002	
				32 350	(o)	3	237.551	8.23E-01	6.90E+08	1.08E+09	-0.24	-0.04	0.251	
				39 115	(o)	1	283.056	2.83E-03	2.37E+06	3.72E+06	-2.56	-2.35	0.244	
				39 346	(o)	2	284.916	1.05E-02	8.83E+06	1.39E+07	-1.98	-1.77	0.660	
				55 715	(o)	1	534.098	4.13E-04	3.46E+05	5.43E+05	-2.85	-2.63	0.000	
				66 048	(o)	2	1192.292	3.94E-02	3.30E+07	5.18E+07	-0.20	0.04	0.693	
				66 390	(o)	1	1242.891	4.01E-05	3.36E+04	5.27E+04	-3.17	-2.91	0.227	
				66 493	(o)	2	1259.000	6.15E-05	5.16E+04	8.10E+04	-2.97	-2.72	0.112	
				66 460	(o)	2	1253.786	1.25E-03	1.05E+06	1.65E+06	-1.66	-1.41	0.403	
				66 564	(o)	3	1270.370	2.48E-04	2.08E+05	3.26E+05	-2.36	-2.10	0.374	
				66 584	(o)	3	1273.628	8.54E-05	7.16E+04	1.12E+05	-2.82	-2.56	0.390	
				67 298	(o)	1	1401.037	5.56E-05	4.66E+04	7.31E+04	-2.93	-2.67	0.085	
				67 396	(o)	2	1420.650	7.89E-07	6.62E+02	1.04E+03	-4.76	-4.50	0.003	
				67 744	(o)	3	1494.454	2.93E-02	2.46E+07	3.86E+07	-0.15	0.11	0.695	
				68 498	(o)	1	1684.392	5.06E-03	4.24E+06	6.65E+06	-0.82	-0.55	0.165	
76 361	(e)	1	1.14 <sup>e</sup>	26 081	(o)	2	198.888	2.24E-04	5.75E+05	5.91E+05	-3.47	-3.46	0.126	
				27 444	(o)	2	204.362	3.68E-04	9.44E+05	9.70E+05	-3.23	-3.22	0.387	
				27 918	(o)	1	206.362	7.57E-02	1.94E+08	1.99E+08	-0.91	-0.90	0.507	
				28 021	(o)	2	206.804	2.30E-01	5.88E+08	6.04E+08	-0.42	-0.41	0.527	
				29 736	(o)	0	214.412	9.25E-05	2.37E+05	2.43E+05	-3.79	-3.77	0.000	
				29 742	(o)	1	214.439	2.71E-04	6.93E+05	7.12E+05	-3.32	-3.31	0.002	
				29 824	(o)	2	214.816	5.31E-05	1.36E+05	1.40E+05	-4.03	-4.01	0.000	
				30 816	(o)	1	219.494	9.17E-05	2.35E+05	2.41E+05	-3.77	-3.76	0.015	
				39 002	(o)	0	267.597	2.33E-01	5.96E+08	6.12E+08	-0.19	-0.18	0.708	
				39 115	(o)	1	268.407	1.73E-01	4.44E+08	4.56E+08	-0.32	-0.31	0.705	
				39 346	(o)	2	270.079	2.84E-01	7.28E+08	7.48E+08	-0.10	-0.09	0.709	
				55 715	(o)	1	484.233	1.12E-06	2.88E+03	2.96E+03	-5.00	-4.98	0.080	
				66 048	(o)	2	969.440	1.10E-05	2.83E+04	2.91E+04	-3.40	-3.39	0.091	
				66 390	(o)	1	1002.628	2.87E-04	7.35E+05	7.55E+05	-1.96	-1.94	0.402	
				66 493	(o)	2	1013.085	4.41E-04	1.13E+06	1.16E+06	-1.76	-1.75	0.319	
				66 460	(o)	2	1009.706	1.72E-04	4.40E+05	4.52E+05	-2.17	-2.16	0.307	
				67 237	(o)	0	1095.698	9.72E-04	2.49E+06	2.56E+06	-1.35	-1.34	0.575	
				67 298	(o)	1	1103.071	6.87E-04	1.76E+06	1.81E+06	-1.49	-1.48	0.528	
				67 396	(o)	2	1115.192	9.76E-04	2.50E+06	2.57E+06	-1.33	-1.32	0.472	
				68 498	(o)	1	1271.473	3.72E-07	9.52E+02	9.78E+02	-4.64	-4.63	0.029	
				75 308	(o)	2	9497.064	1.05E-07	2.69E+02	2.76E+02	-3.44	-3.43	0.567	
				75 591	(o)	2	12 984.147	3.27E-07	8.37E+02	8.60E+02	-2.68	-2.66	0.581	
				75 651	(o)	1	14 082.652	2.29E-07	5.87E+02	6.03E+02	-2.76	-2.75	0.295	
				75 681	(o)	2	14 699.495	9.76E-07	2.50E+03	2.57E+03	-2.10	-2.08	0.578	
75 913	(o)	2	22 304.392	1.26E-08	3.22E+01	3.31E+01	-3.64	-3.61	0.035					
75 952	(o)	1	24 447.995	4.84E-08	1.24E+02	1.27E+02	-2.97	-2.94	0.375					

Table 6 – continued

Upper level <sup>d</sup>	$\tau_u$ (ns)	Lower level <sup>d</sup>	$\lambda^b$ (nm)	$BF$	$gA$ (s <sup>-1</sup> )	$gA_{\text{resc}}$ (s <sup>-1</sup> )	$\log(gf)$	$\log(gf)_{\text{resc}}$	CF <sup>c</sup>
		75 994 (o) 0	27 287.372	2.90E-08	7.44E+01	7.64E+01	-3.11	-3.07	0.363
		76 073 (o) 1	34 770.709	1.85E-10	4.75E-01	4.88E-01	-5.04	-5.05	0.031
76 589 (e) 2	1.09 <sup>e</sup>	26 081 (o) 2	197.989	1.79E-04	7.57E+05	8.23E+05	-3.35	-3.32	0.040
		27 444 (o) 2	203.412	7.58E-05	3.20E+05	3.48E+05	-3.70	-3.67	0.170
		27 602 (o) 3	204.071	1.10E-03	4.64E+06	5.04E+06	-2.54	-2.50	0.490
		27 918 (o) 1	205.393	3.08E-03	1.30E+07	1.41E+07	-2.08	-2.05	0.390
		28 021 (o) 2	205.831	4.50E-02	1.90E+08	2.07E+08	-0.92	-0.88	0.454
		28 161 (o) 3	206.426	2.51E-01	1.06E+09	1.15E+09	-0.17	-0.13	0.529
		29 742 (o) 1	213.393	1.33E-05	5.60E+04	6.09E+04	-4.42	-4.38	0.000
		29 824 (o) 2	213.766	4.74E-04	2.00E+06	2.17E+06	-2.86	-2.83	0.001
		30 816 (o) 1	218.398	1.25E-04	5.27E+05	5.73E+05	-3.42	-3.39	0.012
		32 350 (o) 3	225.973	1.25E-03	5.26E+06	5.72E+06	-2.39	-2.36	0.328
		39 115 (o) 1	266.771	1.77E-01	7.46E+08	8.11E+08	-0.10	-0.06	0.705
		39 346 (o) 2	268.422	5.17E-01	2.18E+09	2.37E+09	0.37	0.41	0.709
		55 715 (o) 1	478.932	2.35E-05	9.91E+04	1.08E+05	-3.47	-3.43	0.019
		66 048 (o) 2	948.425	2.61E-05	1.10E+05	1.20E+05	-2.82	-2.79	0.065
		66 390 (o) 1	980.166	1.23E-05	5.20E+04	5.65E+04	-3.13	-3.09	0.219
		66 493 (o) 2	990.157	1.52E-04	6.43E+05	6.99E+05	-2.03	-1.99	0.337
		66 460 (o) 2	986.929	6.09E-05	2.57E+05	2.79E+05	-2.42	-2.39	0.243
		66 564 (o) 3	997.176	6.97E-04	2.94E+06	3.20E+06	-1.36	-1.32	0.372
		66 584 (o) 3	999.182	1.61E-04	6.78E+05	7.37E+05	-1.99	-1.96	0.215
		67 298 (o) 1	1075.944	8.96E-04	3.78E+06	4.11E+06	-1.18	-1.15	0.593
		67 396 (o) 2	1087.473	2.06E-03	8.71E+06	9.47E+06	-0.81	-0.77	0.517
		67 744 (o) 3	1130.199	9.83E-05	4.15E+05	4.51E+05	-2.10	-2.06	0.549
		68 498 (o) 1	1235.566	1.70E-05	7.17E+04	7.79E+04	-2.78	-2.75	0.138
		75 308 (o) 2	7803.238	4.76E-08	2.01E+02	2.18E+02	-3.74	-3.70	0.536
		75 309 (o) 3	7806.711	2.44E-07	1.03E+03	1.12E+03	-3.03	-2.99	0.670
		75 373 (o) 3	8220.091	2.89E-07	1.22E+03	1.33E+03	-2.91	-2.87	0.502
		75 552 (o) 3	9642.061	1.18E-07	5.00E+02	5.43E+02	-3.16	-3.12	0.304
		75 591 (o) 2	10 012.694	8.81E-09	3.72E+01	4.04E+01	-4.26	-4.22	0.012
		75 651 (o) 1	10 653.532	2.30E-09	9.70E+00	1.05E+01	-4.79	-4.75	0.011
		75 681 (o) 2	11 002.823	1.28E-07	5.40E+02	5.87E+02	-3.02	-2.97	0.092
		75 716 (o) 3	11 444.422	2.21E-06	9.33E+03	1.01E+04	-1.75	-1.70	0.611
		75 913 (o) 2	14 773.133	5.78E-07	2.44E+03	2.65E+03	-2.11	-2.06	0.445
		75 952 (o) 1	15 683.966	1.11E-07	4.69E+02	5.10E+02	-2.78	-2.73	0.354
77 387 (e) 3	3.73 <sup>e</sup>	76 073 (o) 1	19 373.811	7.70E-09	3.25E+01	3.53E+01	-3.73	-3.70	0.278
		26 081 (o) 2	194.910	2.66E-05	3.37E+04	5.00E+04	-4.71	-4.55	0.015
		27 444 (o) 2	200.162	8.78E-04	1.11E+06	1.65E+06	-3.17	-3.00	0.300
		27 602 (o) 3	200.800	2.37E-02	3.00E+07	4.45E+07	-1.74	-1.57	0.312
		27 841 (o) 4	201.768	2.02E-01	2.56E+08	3.80E+08	-0.80	-0.63	0.327
		28 021 (o) 2	202.504	1.99E-02	2.52E+07	3.74E+07	-1.81	-1.64	0.293
		28 161 (o) 3	203.079	1.29E-01	1.63E+08	2.42E+08	-0.99	-0.82	0.269
		29 824 (o) 2	210.180	1.76E-01	2.23E+08	3.31E+08	-0.83	-0.66	0.374
		32 350 (o) 3	221.970	4.87E-05	6.16E+04	9.14E+04	-4.34	-4.17	0.126
		39 346 (o) 2	262.791	1.16E-01	1.47E+08	2.18E+08	-0.81	-0.65	0.249
		66 048 (o) 2	881.687	2.62E-04	3.32E+05	4.93E+05	-2.39	-2.24	0.123
		66 493 (o) 2	917.642	8.86E-03	1.12E+07	1.66E+07	-0.83	-0.68	0.509
		66 460 (o) 2	914.869	6.78E-03	8.58E+06	1.27E+07	-0.95	-0.80	0.781
		66 564 (o) 3	923.667	6.02E-02	7.62E+07	1.13E+08	0.01	0.16	0.379
		66 584 (o) 3	925.388	7.61E-02	9.63E+07	1.43E+08	0.11	0.26	0.792
		66 719 (o) 4	937.110	1.33E-01	1.68E+08	2.49E+08	0.36	0.52	0.748
		67 396 (o) 2	1000.629	4.64E-02	5.87E+07	8.71E+07	-0.04	0.12	0.575
		67 744 (o) 3	1036.689	4.70E-05	5.94E+04	8.81E+04	-3.00	-2.85	0.115
		75 221 (o) 4	4616.186	1.04E-06	1.31E+03	1.94E+03	-3.28	-3.21	0.152
		75 308 (o) 2	4808.601	2.30E-08	2.91E+01	4.32E+01	-4.90	-4.82	0.068
		75 309 (o) 3	4809.920	4.80E-08	6.07E+01	9.01E+01	-4.58	-4.50	0.011
		75 373 (o) 3	4963.717	4.93E-07	6.24E+02	9.26E+02	-3.54	-3.47	0.150
		75 390 (o) 4	5006.948	1.42E-06	1.80E+03	2.67E+03	-3.07	-3.00	0.104
		75 470 (o) 4	5215.252	1.24E-06	1.57E+03	2.33E+03	-3.09	-3.02	0.112
		75 552 (o) 3	5448.967	1.03E-06	1.30E+03	1.93E+03	-3.13	-3.07	0.538
		75 561 (o) 4	5474.449	1.18E-08	1.49E+01	2.21E+01	-5.05	-5.00	0.107
		75 591 (o) 2	5565.388	5.53E-07	7.00E+02	1.04E+03	-3.37	-3.32	0.306

Table 6 – continued

Upper level <sup>d</sup>	$\tau_{ii}$ (ns)	Lower level <sup>d</sup>	$\lambda^b$ (nm)	$BF$	$gA$ (s <sup>-1</sup> )	$gA_{\text{resc}}$ (s <sup>-1</sup> )	$\log(gf)$	$\log(gf)_{\text{resc}}$	$CF^c$
75 681	(o)	2	5858.418	1.88E-06	2.38E+03	3.53E+03	-2.79	-2.74	0.571
75 716	(o)	3	5981.306	1.01E-05	1.28E+04	1.90E+04	-2.04	-1.99	0.848
75 913	(o)	2	6779.698	4.40E-07	5.56E+02	8.25E+02	-3.28	-3.24	0.072

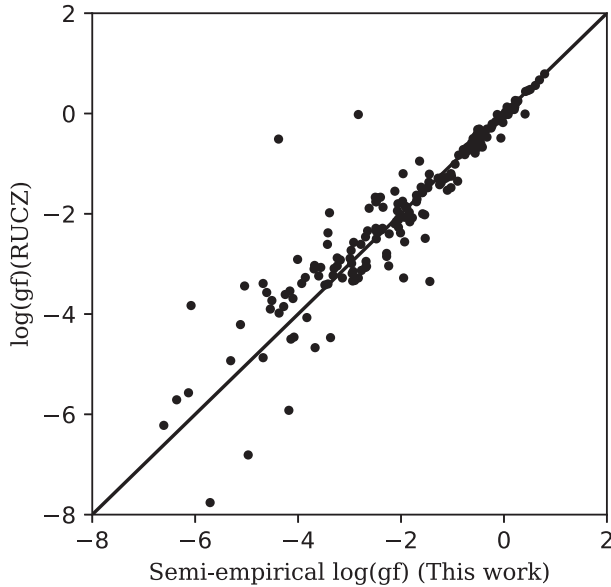
Notes. <sup>a</sup>Each level is designated by its value in cm<sup>-1</sup>, its parity ((e) and (o) stand for even and odd, respectively) and its total quantum number,  $J$ .

<sup>b</sup>Calculated from the energy levels compiled by NIST (Kramida et al. 2015).  $\lambda > 200$  nm are given in air.

<sup>c</sup>Cancellation factor as defined by Cowan (1981). A value less than 5 per cent indicates a strong cancellation effect on the line strength and the transition probability could be underestimated.

<sup>d</sup>TR-LIF measurements by Marsden et al. (1988).

<sup>e</sup>TR-LIF measurements, this work.

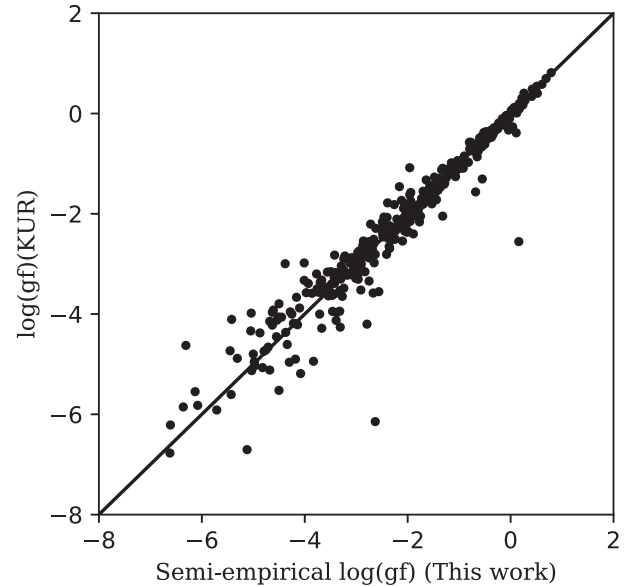


**Figure 4.** A comparison between the oscillator strengths determined by the combination of the HFR+CPOL branching fractions and the TR-LIF lifetimes (this work) and the semi-empirical oscillator strengths calculated by Ruczkowski et al. (2014) (RUCZ). The straight line of equality has been drawn.

the meteoritic abundances (Lodders et al. 2009) still remains for scandium with  $\log \epsilon_{\odot} = 3.16 \pm 0.04$  and  $\log \epsilon_{\text{met}} = 3.05 \pm 0.02$ .

The Sc II lines used in Scott et al. (2015) for the determination of the solar abundance of scandium are presented in Table 7. These lines are from low-excited levels measured by Lawler & Dakin (1989) but not included in this study. The third column of this table contains the oscillator strengths deduced from the  $A$ -values of Lawler & Dakin (1989) used by Scott et al. (2015) to determine the photospheric abundances listed in the sixth column. They are compared to our rescaled oscillator strengths reported in the fourth column and the differences between the two values are given in the log scale in the fifth column. For this set of solar lines, our oscillator strengths are systematically larger than those of Lawler & Dakin (1989) by  $\sim 0.1$  dex on average, if we exclude the transition  $3d^2 \ ^1D_2 - 3d4p \ ^1D_2^{\circ}$  for which our  $f$ -value is affected by strong cancellation effects. Column seven in Table 7 gives the abundances obtained from each line, assuming they are all lying on the linear part of the curve of growth (see the upper left panel of fig. 3 in Scott et al. (2015)), with our new  $gf$ -values.

The weighted average along with the corresponding weighted standard deviation of the abundance was determined using the weights of Scott et al. (2015), reported in the last column of



**Figure 5.** A comparison between the oscillator strengths determined by the combination of the HFR+CPOL branching fractions and the TR-LIF lifetimes (this work) and the oscillator strengths calculated by Kurucz (2011) (KUR). The straight line of equality has been drawn.

Table 7. Their weights range from one to three and are based on the line quality for abundance determination. Discarding the line  $3d^2 \ ^1D_2 - 3d4p \ ^1D_2^{\circ}$  from the mean estimate, one obtains  $\log \epsilon_{\text{cor}} = 3.04 \pm 0.13$  (where the second number is the standard deviation) for the corrected photospheric abundance, now in good agreement with the meteoritic value of Lodders et al. (2009). Even if we reject the transition  $3d^2 \ ^3F_4 - 3d4p \ ^3F_3^{\circ}$  for which there is a factor of two difference between our rescaled  $f$ -value and the experimental value of Lawler & Dakin (1989), the mean  $\log \epsilon_{\text{cor}} = 3.10 \pm 0.05$  is still in accord with the meteoritic value. Moreover, considering the full line set does not change the agreement ( $\log \epsilon_{\text{cor}} = 3.07 \pm 0.17$ ). Finally, we note that all these weighted average abundances agree within the mutual error bars with the value determined by Scott et al. (2015) using only Sc I lines ( $\log \epsilon = 3.14 \pm 0.09$ ).

Replacing our  $f$ -value set by the one of Kurucz (2011) will not change this accord either ( $\log \epsilon_{\text{kur}} = 3.10 \pm 0.09$ ). This is not the case for the set of Ruczkowski et al. (2014). Indeed, the photospheric abundance would be estimated significantly too high with respect to the meteoritic value, i.e.  $\log \epsilon_{\text{ruc}} = 3.44 \pm 0.31$ . Even if the transition  $3d^2 \ ^3F_4 - 3d4p \ ^3F_3^{\circ}$  for which the oscillator strength calculated by Ruczkowski et al. (2014) ( $\log(gf)_{\text{ruc}} = -3.28$ ) is one order of magnitude lower than the experimental value of Lawler &

**Table 7.** Sc II lines used in the determination of the solar abundance ( $\log \epsilon$ ) of scandium.

Transition		$\lambda^a$ (nm)	$\log(gf)_{L\&D}^b$	$\log(gf)_{resc}^c$	$\Delta \log(gf)^d$	$\log \epsilon^e$	$\log \epsilon_{cor}^f$	Weight <sup>e</sup>	
$3d^2\ ^3F_4$	–	$3d4p\ ^3F_3^o$	442.067	–2.273	–1.950	0.323	3.099	2.776	2
$3d^2\ ^3F_3$	–	$3d4p\ ^3F_2^o$	443.135	–1.969	–1.780	0.189	3.155	2.966	1
$3d^2\ ^3P_2$	–	$3d4p\ ^1P_1^o$	535.720	–2.111	–2.050	0.061	3.131	3.070	2
$3d^2\ ^3P_1$	–	$3d4p\ ^3P_2^o$	564.100	–1.131	–1.000	0.131	3.226	3.095	1
$3d^2\ ^3P_0$	–	$3d4p\ ^3P_1^o$	565.836	–1.208	–1.150	0.058	3.211	3.153	1
$3d^2\ ^3P_1$	–	$3d4p\ ^3P_1^o$	566.715	–1.309	–1.220	0.089	3.235	3.146	1
$3d^2\ ^3P_1$	–	$3d4p\ ^3P_0^o$	566.904	–1.200	–1.100	0.100	3.246	3.146	1
$3d^2\ ^3P_2$	–	$3d4p\ ^3P_1^o$	568.420	–1.074	–1.030	0.044	3.154	3.110	2
$3d^2\ ^1D_2$	–	$3d4p\ ^1D_2^o$	660.460	–1.309	–1.570 <sup>g</sup>	–0.261	3.204	3.465	1

Notes. <sup>a</sup>Calculated from the energy levels compiled by NIST (Kramida et al. 2015).

<sup>b</sup>Scott et al. (2015), deduced from the  $A$ -values of Lawler & Dakin (1989).

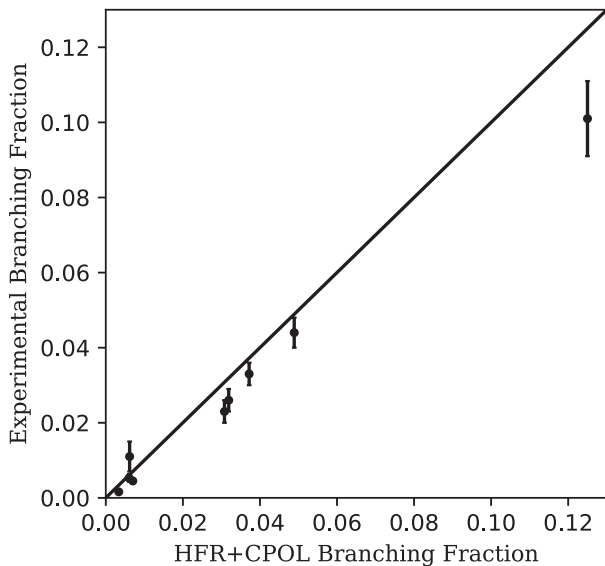
<sup>c</sup>This work, rescaled value.

<sup>d</sup> $\Delta \log(gf) = \log(gf)_{L\&D} - \log(gf)_{resc}$ .

<sup>e</sup>Scott et al. (2015).

<sup>f</sup>This work, corrected abundance,  $\log \epsilon_{cor} = \log \epsilon - \Delta \log(gf)$ .

<sup>g</sup>Affected by strong cancellation effects ( $CF < 0.05$ ).



**Figure 6.** A comparison between the HFR+CPOL branching fractions and the experimental values of Lawler & Dakin (1989) for the Sc II lines used in the determination of the solar scandium abundance. The straight line of equality has been drawn.

Dakin (1989) is excluded, this would not significantly improve the situation ( $\log \epsilon_{resc} = 3.29 \pm 0.01$ ).

It should be noted, however, that the lines used for these studies are weak, see Fig. 6. Their  $BF$ s are less than 5 per cent except for  $\lambda 566.904$  having  $\sim 10$  per cent. These small  $BF$ s make it difficult to measure and calculate with high accuracy. The real uncertainty might thus be larger than the observed scatter.

## 7 CONCLUSIONS

New TR-LIF lifetimes were measured using two-step excitation schemes in Sc II. These measurements extend the set of available experimental values with 17 even-parity levels belonging to the excited configurations  $3d5s$ ,  $3d4d$ ,  $4p^2$  and  $3d6s$ . We measured 57  $BF$ s from these upper levels using an HCL and an FTS. By combining the  $BF$ s with the measured lifetimes, we derived  $\log(gf)$  values from these highly excited levels. A HFR model that includes

the main relativistic interactions along with the core-polarization effects (HFR+CPOL) was used to determine the branching fractions and the oscillator strengths. The comparison between our HFR+CPOL and TR-LIF lifetimes along with those found in the literature (Buchta et al. 1971; Arnesen et al. 1976; Palenius et al. 1976; Vogel et al. 1985; Marsden et al. 1988; Kurucz 2011; Ruczkowski et al. 2014) shows generally a good agreement ranging from a few per cent to 20 per cent with the notable exceptions of the even-parity levels  $4p^2\ ^1D_2$  and  $3d6s\ ^3D_3$ . The former discrepancy may be due to a cancellation effect that lengthens the HFR+CPOL lifetime. Owing to the good agreement ( $\sim 20$  per cent) obtained with the experimental branching fractions of Lawler & Dakin (1989) for low-excitation levels and ours for high-excitation levels, the HFR+CPOL branching fractions were combined with our TR-LIF lifetimes and the experimental values of Marsden et al. (1988) to obtain rescaled semi-empirical oscillator strengths for all the 380 E1 transitions depopulating the 34 fine-structure levels for which TR-LIF lifetimes are available. This new set of oscillator strengths was compared to the parametric calculation of Ruczkowski et al. (2014) and to the HFR values of Kurucz (2011). In both cases, the mean scatters were  $\sim 20$  per cent for the strong lines ( $\log(gf) \gtrsim -1$ ) giving an estimate of the accuracy for these radiative parameters. Finally, the solar abundance of scandium was estimated to  $\log \epsilon_{\odot} = 3.04 \pm 0.13$  using our rescaled semi-empirical oscillator strengths to correct the values determined in the recent study of Scott et al. (2015). This value is in improved agreement with the meteoritic value ( $\log \epsilon_{met} = 3.05 \pm 0.02$ ) of Lodders et al. (2009).

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