

Pollution and climate modelling

An Introduction

Sources: “*Fundamentals of Atmospheric Modelling*”, M. Jacobson (2005); “*Chemical Transport Models*”, online book
by D. Jacob (<http://acmg.seas.harvard.edu/education/>)

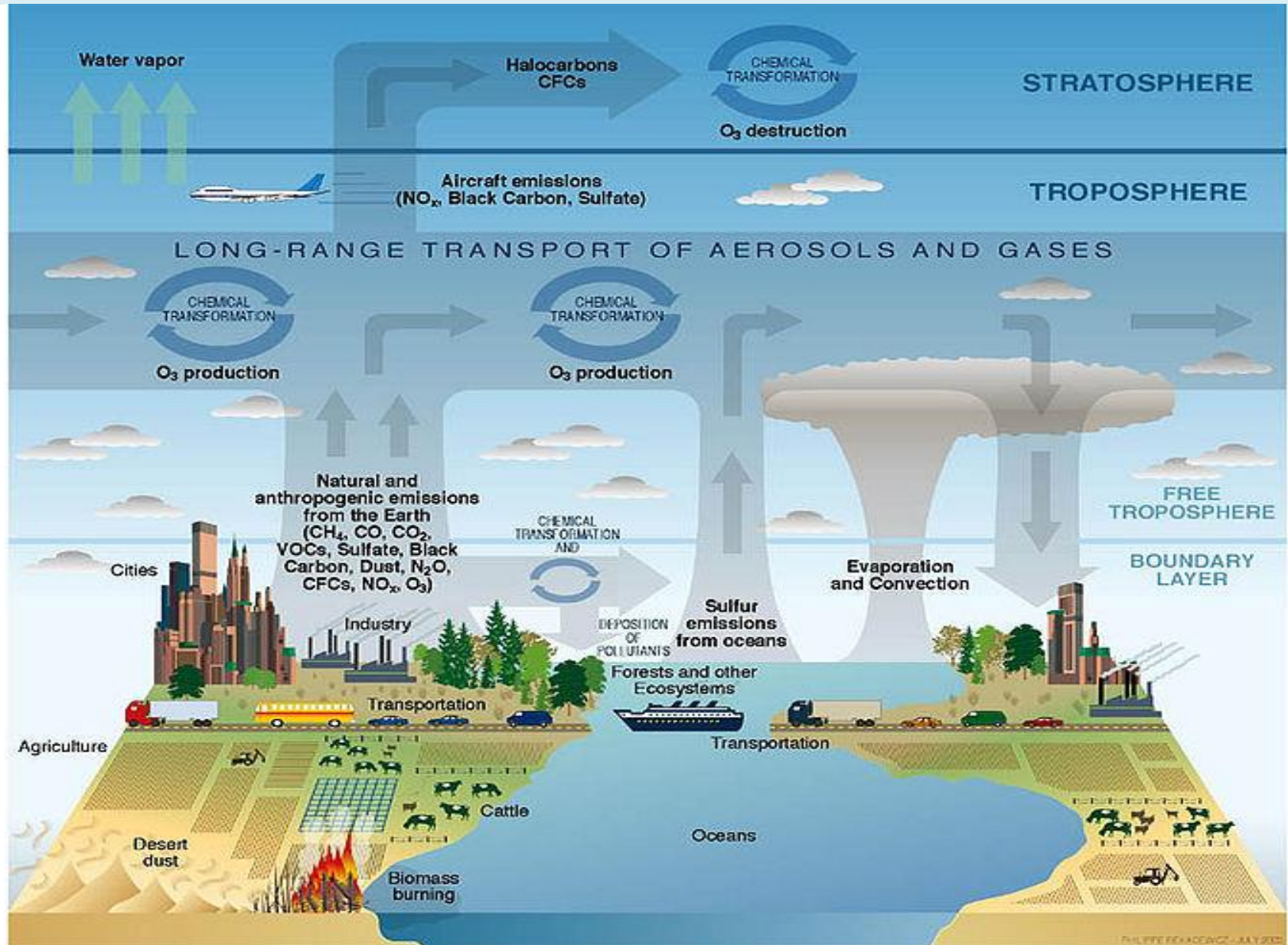
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SPAT PG Lectures, 1st of December 2016

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The (complex) composition-climate system:



Source: US Climate Change Science Program.

Recent advances

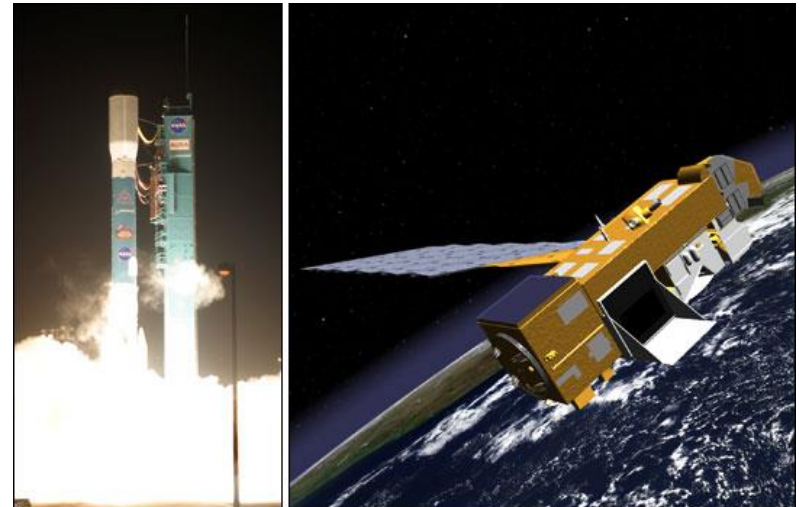
Models: • **Chemistry-climate models** are climate models with atmospheric composition (gases, aerosols) “on top”.

- They have advanced a lot in the last 2-3 decades, but they can improve even more.

Satellites: **Observations of atmospheric constituents** have produced a wealth of data (e.g. NASA A-Train), especially in the last decade.



Discover supercomputer



Boeing Delta II Rocket

Artist's rendering of Aura in Orbit

Aura satellite

Mass balance (continuity) equation for a gas/aerosol constituent

$$\frac{\partial n}{\partial t} = -\nabla \cdot (n\mathbf{U}) + D\nabla^2 n + P - L$$

Change of concentration of constituent with time (in molec. $\text{m}^{-3} \text{s}^{-1}$).

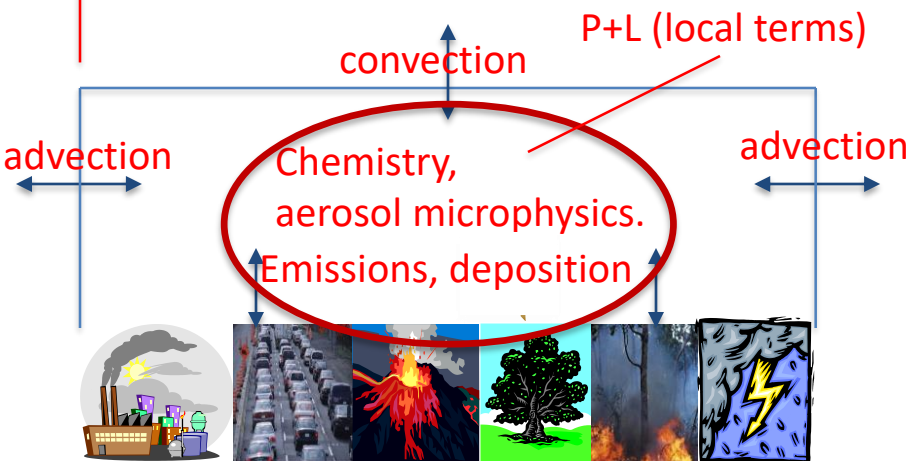
Flux divergence due to transport (advection/convection). \mathbf{U} is the wind velocity vector (m s^{-1}).

Flux divergence due to molecular diffusion. D is the molecular diffusion coefficient ($\text{m}^2 \text{s}^{-1}$).

Local production term: Emission, chemical production, microphysics (in $\text{kg m}^{-3} \text{s}^{-1}$).

Local loss term: Chemical loss, wet and dry deposition, microphysics (in $\text{kg m}^{-3} \text{s}^{-1}$).

Small in trop/strat

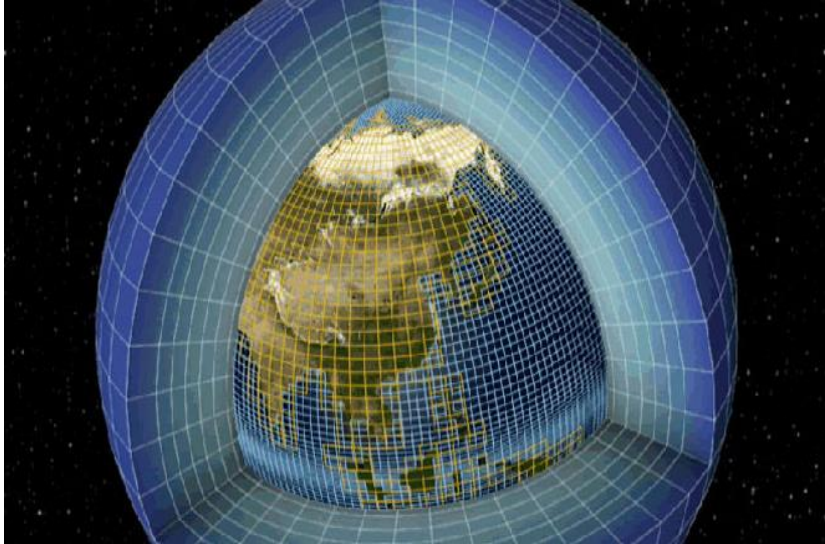


- This equation is for number concentration, but equivalent equations can be written for **mass**, **mass concentration** etc.

Discretisation of mass balance equation in space

(useful for global modelling)

- First order PDE in space and time. Need **initial concentration** and **boundary conditions** (i.e. fluxes at surface, top-of-the-atmosphere).
- Global models have 3D domain with finite number of **gridboxes**. Typical global models: horizontal resolution of $\sim 100\text{km}$, vertical of $\sim 1\text{km}$ \rightarrow total of $\sim 10^6$ gridboxes. **Equation then solved for all gridboxes.**



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- **Chemical transport models** (CTMs): use external meteorological data as input; simulate the aerosols/gases.
 - **General circulation models** (GCMs): simulate their own meteorology; use external aerosols gas forcings.
- Composition-climate models** (CCMs): **do both.**

Discretisation of mass balance equation in time

- **Split the equation** into contributions from transport and local terms:

$$\frac{\partial n}{\partial t} = \left[\frac{\partial n}{\partial t} \right]_{TRANSPORT} + \left[\frac{\partial n}{\partial t} \right]_{LOCAL}$$

Advection+convection = $-\nabla \cdot (n\mathbf{U})$

Emissions, chemistry, deposition etc ($P-L$).
If several species, then for species i we have $P_i(\mathbf{n})-L_i(\mathbf{n})$, where \mathbf{n} a vector of concentrations of all species on which i depends.

- Use a **transport operator** and a **local operator** to **decouple** the two terms in finite difference form (assumed TRA and LOC are decoupled - can swap to test!):

$$n(t_0 + \Delta t) = (\mathbf{LOC}) \cdot (\mathbf{TRA})(n(t_0))$$

- Can **split further** (TRA to TRA_x , TRA_y , TRA_z , or LOC to chemistry, microphysics, emissions, deposition operators).

Photolysis rate calculation

- For a given molecule A being photolyzed ($A+h\nu \longrightarrow B+C$):

$$J_A(T) = \int \sigma_A(\lambda) \varphi_A(\lambda) F(\lambda) d\lambda \quad (s^{-1})$$

- $\sigma_A(\lambda)$: Absorption **cross section** (probability of a photon to be absorbed).
- $\varphi_A(\lambda)$: **Quantum yield** (number of molecules photolyzed per photon absorbed).
- $F(\lambda)$: Solar **actinic flux** (radiative flux from all directions).
- When photolysis rate multiplied by concentration -> gives **loss rate** of constituent (mass loss per unit time).
- Generally rate of reaction: $r = k(T)[A]^m[B]^n$, (k =reaction rate constant).

Lifetimes of constituents

- The rate at which a chemical species A (with concentration $[A]$) is lost from the atmosphere is characterised by **its e-folding lifetime** τ_A .
- It is the time required by a gas to decrease to $1/e$ its original concentration due to chemical reaction. Lifetimes are **independent** of emission/production rates.

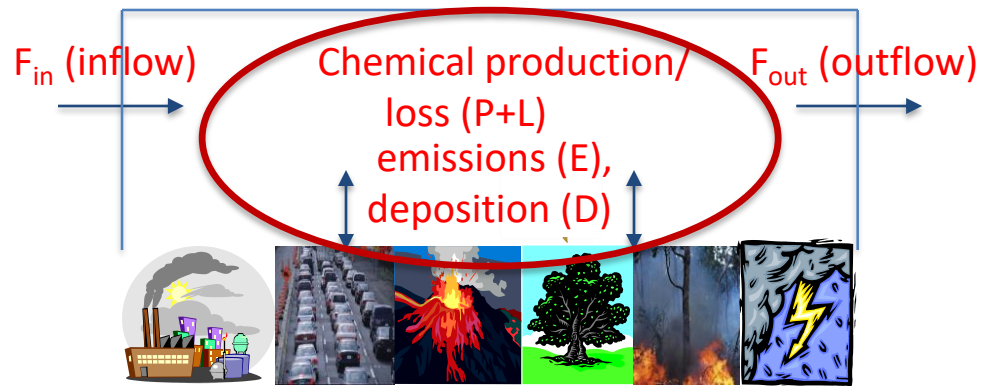
- For example, for **photolysis** reaction:
$$\tau_A = \frac{1}{J_A} = \frac{[A]}{d[A]/dt}$$

- If concentrations of a species are determined by multiple processes (1, 2, ..., n), the overall lifetime is:

$$\tau_A = \frac{1}{\frac{1}{\tau_{A1}} + \frac{1}{\tau_{A2}} + \dots + \frac{1}{\tau_{An}}}$$

Simple one-box models (*0-dimensional*)

- Helpful for quickly testing hypotheses. Also OK for modelling WMGHGs (not very inhomogeneous in space).
- Consider gas (no microphysics) with mass m :



$$\frac{\partial m}{\partial t} = \sum \text{sources} - \sum \text{sinks} = F_{in} + E + P - F_{out} - L - D$$

$$\text{Overall lifetime: } \tau = \frac{m}{\sum \text{sinks}} = \frac{m}{F_{out} + L + D}$$

Overall loss rate constant (in s^{-1}): $k = 1/\tau = k_{out} + k_L + k_D$ (note: if the only loss process is photolysis, then $k_L = J$).

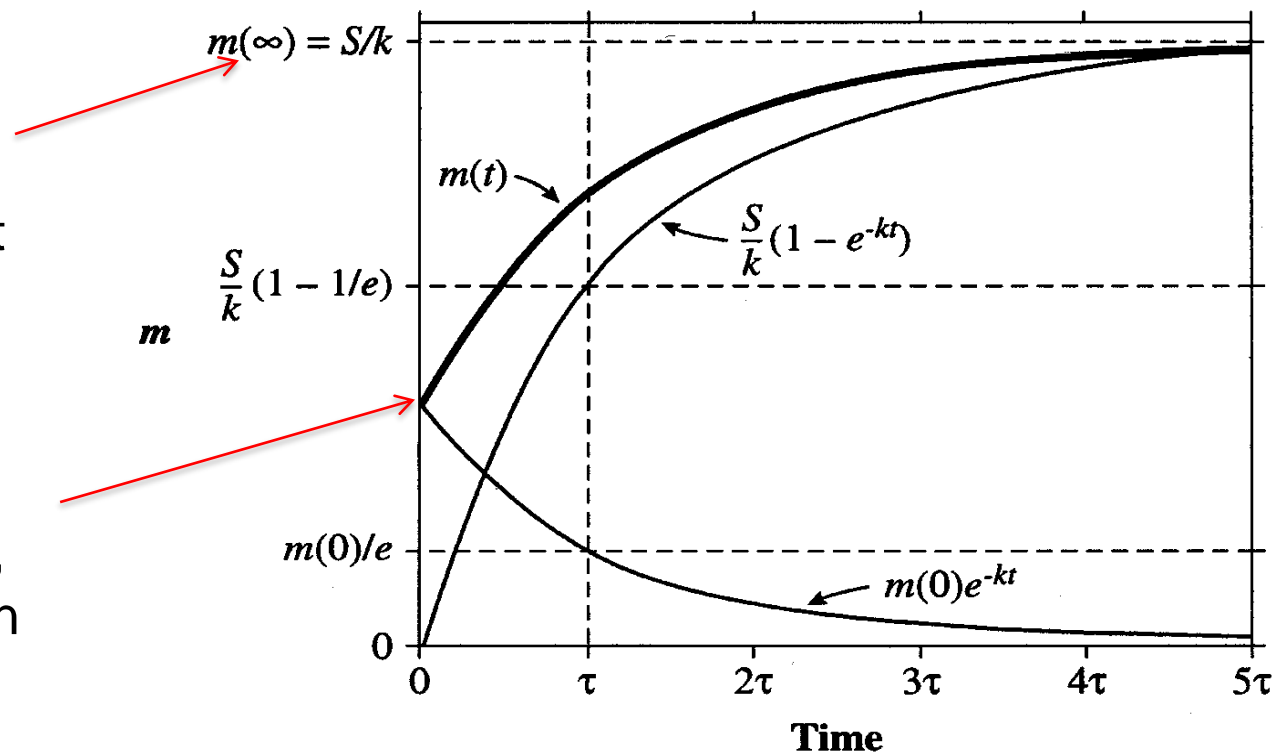
Evolution of constituent mass in a box model

- Assume we have a constant known source S (emission, chemical, microphysical) and constant 1st order loss L (loss rate constant = k).

$$\frac{dm}{dt} = S - L = S - km \quad \Rightarrow \quad m(t) = m(0)e^{-kt} + \frac{S}{k}(1 - e^{-kt})$$

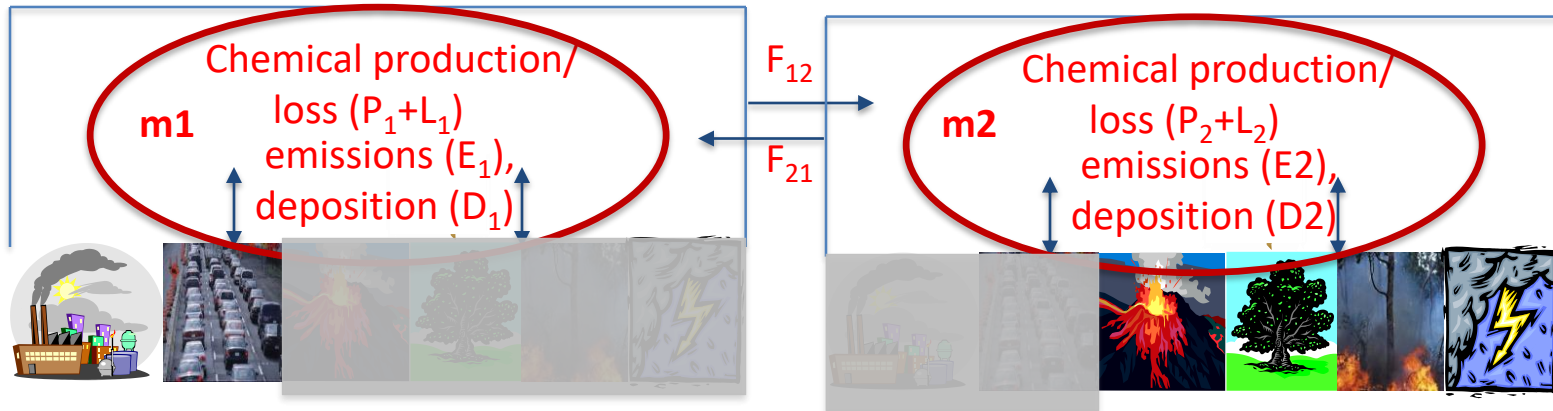
- **Steady state** when $dm/dt = 0$, i.e. abundance does not change with time.

- Takes about 2-3 lifetimes (τ) to reach “**quasi steady-state**”, though it depends on the **initial conditions**.



Simple two-box model

- Useful if we have two distinct environments interacting (e.g. boundary layer & free troposphere, Europe and the Sahara etc).



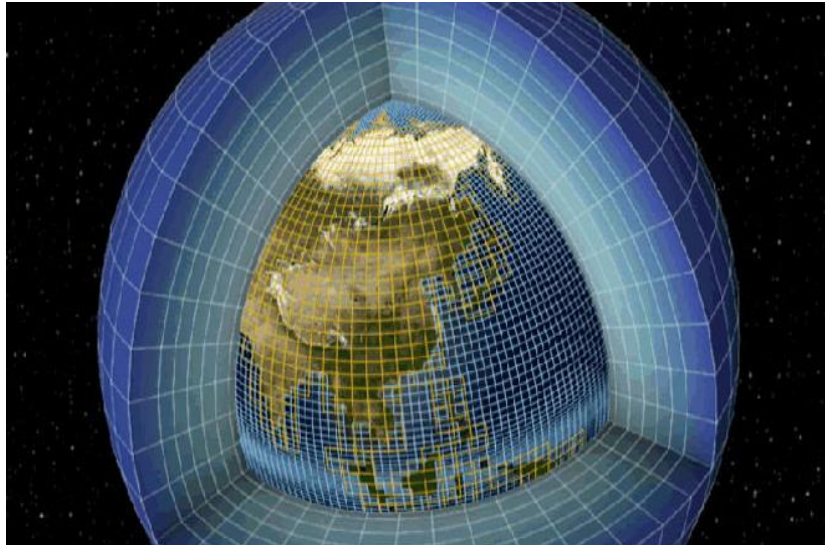
$$\frac{dm_1}{dt} = E_1 + P_1 - L_1 - D_1 - F_{12} + F_{21} \quad (\text{and similar for box 2})$$

We may be given that some loss process is first order (e.g. the left to right transport flux). Then we can express $F_{12} = k_{F12}m_1$.

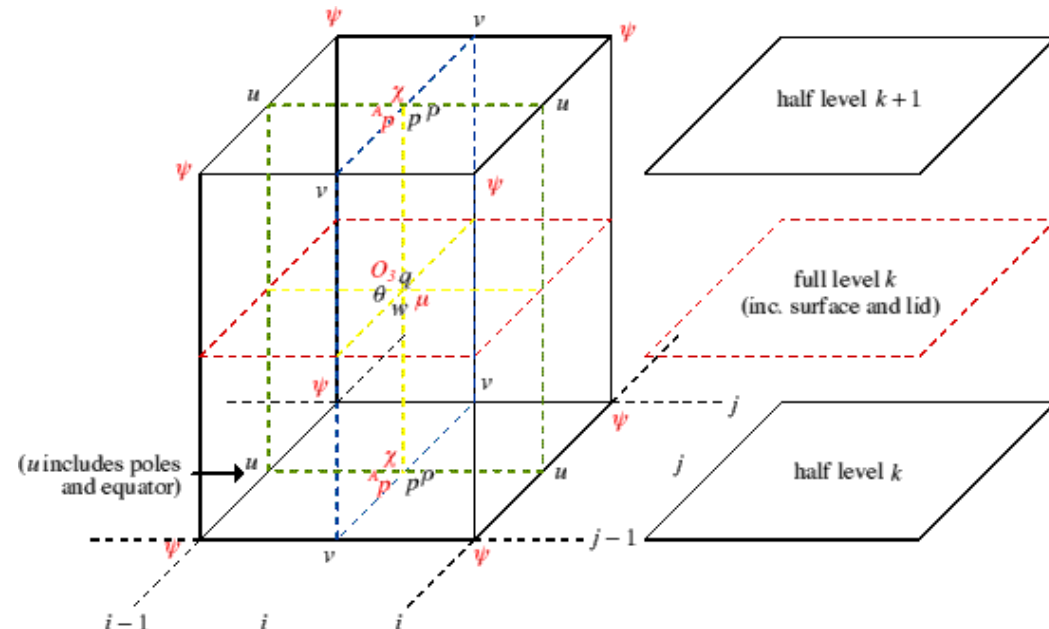
Note that if we are told that the system has reached steady state, we have two simple algebraic equations.

Back to large-scale models

- First order PDE in space and time. Need **initial concentration** and **boundary conditions** (i.e. fluxes at surface, top-of-the-atmosphere).
- Global models have 3D domain with finite number of **gridboxes**. Typical global models: horizontal resolution of $\sim 100\text{km}$, vertical of $\sim 1\text{km}$ \rightarrow total of $\sim 10^6$ gridboxes. **Equation then solved for all gridboxes.**



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<http://www.met.reading.ac.uk/~ross/DARC/Grids.html>

Back to large-scale models: Thermodynamic energy equation

$$\frac{\partial E}{\partial t} = -\nabla \cdot (\mathbf{U}E) + \rho_A \frac{\theta_V}{T_V} \frac{dQ}{dt}$$

Change of energy density with time in a gridbox (in $\text{J m}^{-3} \text{s}^{-1}$).

Flux divergence for energy density due to transport. \mathbf{U} is the wind velocity vector (m s^{-1}).

Heat production and loss terms (ρ_A is the density of air, and θ_V/T_V are the virtual potential temperature and virtual temperature).
Includes diabatic heating/cooling by greenhouse gases/aerosols.

- It is also solved for **each timestep** in a GCM or a CCM. In the CCM **“online” simulated** gases/aerosols will be used for diabatic heating
- Simulated heating rates and temperatures are then **fed back** to the chemistry/aerosol “scheme” to influence atmospheric composition.

Conservation of momentum (*Navier Stokes equation*)

$$\frac{\partial \mathbf{U}}{\partial t} = -(\mathbf{U} \cdot \nabla) \mathbf{U} + \nu \nabla^2 \mathbf{U} - \frac{\nabla p}{\rho_A} + g$$

Change in air velocity in a gridbox (in m s^{-2}). Advection term for \mathbf{U} . Changes due to diffusion. Pressure-gradient & gravity terms, exerting local accelerations.

- Familiar? Analogous to the mass balance equation.
- Heating/cooling rates calculated from thermodynamic balance (previous) affect pressures, which are **fed into this part of the model (dynamics)** to calculate **changes in wind velocities** (i.e. circulation).
- Then circulation changes, together with other resulting meteorological changes that occur during the timestep, **feed into the chemistry/aerosol** scheme to affect constituent transport, reaction rates, wet deposition etc.
- **Everything coupled** (in a CCM)!

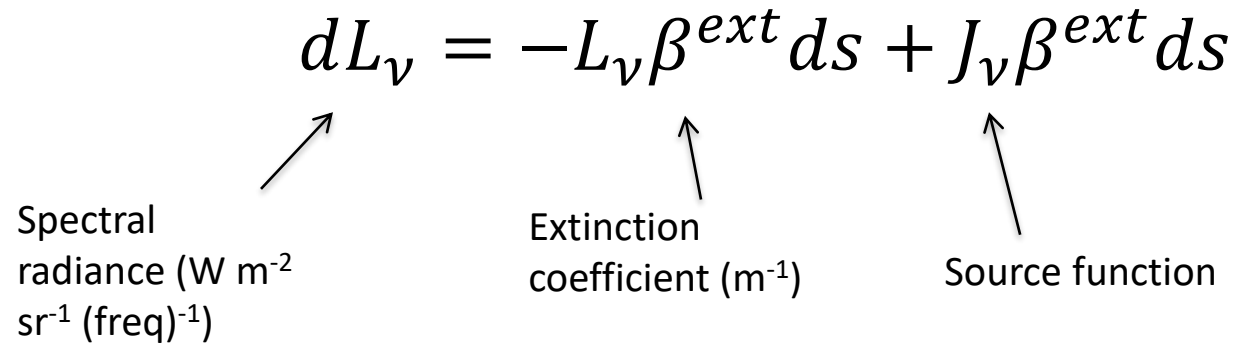
Radiative transfer

$$dL_\nu = -L_\nu \beta^{ext} ds + J_\nu \beta^{ext} ds$$

Spectral radiance ($\text{W m}^{-2} \text{sr}^{-1} (\text{freq})^{-1}$)

Extinction coefficient (m^{-1})

Source function

The diagram shows the radiative transfer equation $dL_\nu = -L_\nu \beta^{ext} ds + J_\nu \beta^{ext} ds$. Three arrows point from labels below to terms in the equation: one from 'Spectral radiance' to dL_ν , one from 'Extinction coefficient' to β^{ext} in the first term, and one from 'Source function' to J_ν in the second term.

- Also needs to be discretized – in **frequency** and **direction** as well as space and time
- One of the source terms (dQ) in thermodynamic equation – typically represents one of the most computationally intensive components of a climate model

Gas processes

Emission
 Photochemistry
 Heterogeneous chemistry
 Aerosol nucleation
 Condensation/evaporation
 Dissolution/evaporation
 Dry deposition
 Washout

Radiative processes

Solar and infrared radiation
 Gas, aerosol, cloud absorption
 Gas, aerosol, cloud scattering
 Heating rates
 Actinic fluxes
 Visibility
 Albedo

Aerosol processes

Emission
 Nucleation
 Aerosol–aerosol coagulation
 Aerosol–hydrometeor coagulation
 Condensation/evaporation
 Dissolution/evaporation
 Equilibrium chemistry
 Aqueous chemistry
 Heterogeneous chemistry
 Dry deposition/sedimentation
 Rainout/washout

Meteorological processes

Air temperature
 Air density
 Air pressure
 Wind speed and direction
 Turbulence
 Water vapor

Cloud processes

Condensation/ice deposition
 Homogeneous, contact freezing
 Melting/evaporation/sublimation
 Hydrometeor–hydrometeor coag.
 Aerosol–hydrometeor coagulation
 Gas dissolution/aqueous chemistry
 Precipitation, rainout, washout
 Lightning

Transport processes

Emission
 Gas, aerosol, cloud transport in air
 Gas, aerosol transport in clouds
 Dry deposition/sedimentation
 Rainout/washout

Surface processes

Soil, water, sea ice, snow, road,
 roof, vegetation temperatures
 Surface energy, moisture fluxes
 Ocean dynamics

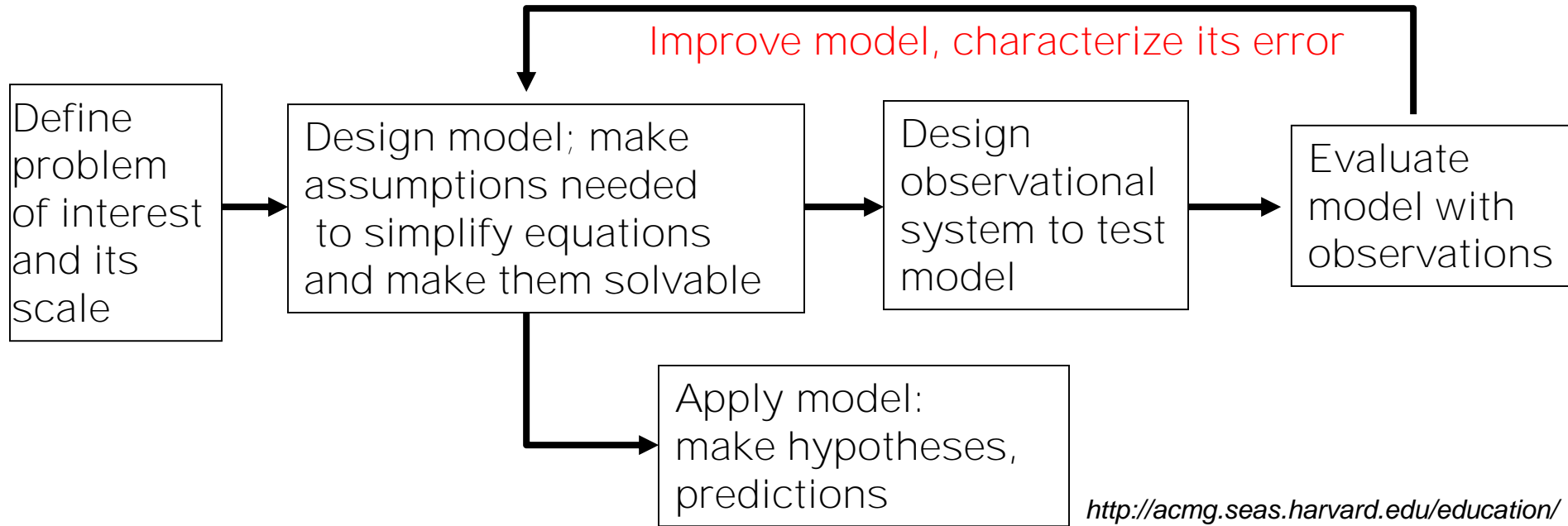
Atmospheric processes simulated in a CCM

Surface processes not strictly atmospheric, but can influence atmosphere. In **Earth system models**, the biosphere is simulated simultaneously, i.e. vegetation type, growth, and emissions **depend on simulated climate**.

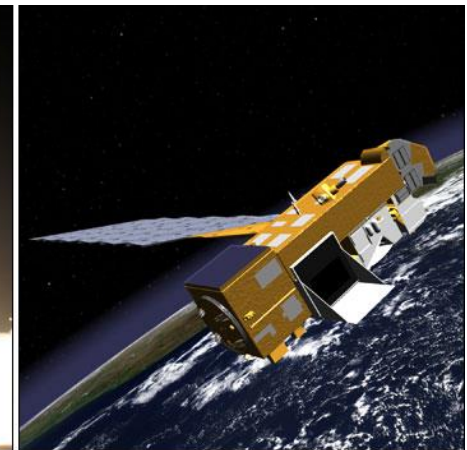
swn2o_toa	SW TOA N2O RADIATIVE FORCING	[lon][lat]
swn2o_toa_hemis	swn2o_toa_hemis	—
swn_grnd_clrsky	CLR SKY NET SOLAR RADIATION, SRF	[lon][lat]
swn_grnd_clrsky_hemis	swn_grnd_clrsky_hemis	—
swncls	SW CLR-SKY NET RADIATION, SURFA...	[lon][lat]
swncls_hemis	swncls_hemis	—
swnclt	SW CLR-SKY NET RADIATION TOA ME...	[lon][lat]
swnclt_hemis	swnclt_hemis	—
swu_oice	SEA ICE UPWARD SHORTWAVE RADIA...	[lon][lat]
swu_oice_hemis	swu_oice_hemis	—
swup_toa_clrsky	CLR SKY OUT SOLAR RADIATION, TOA	[lon][lat]
swup_toa_clrsky_hemis	swup_toa_clrsky_hemis	—
t_300	TEMPERATURE AT 300mb	[lon][lat]
t_300_hemis	t_300_hemis	—
t_500	TEMPERATURE AT 500mb	[lon][lat]
t_500_hemis	t_500_hemis	—
t_700	TEMPERATURE AT 700mb	[lon][lat]
t_700_hemis	t_700_hemis	—
t_850	TEMPERATURE AT 850mb	[lon][lat]
t_850_hemis	t_850_hemis	—
Tatm	ATMOSPHERIC TEMPERATURE	[lon][lat]
Tatm_hemis	Tatm_hemis	—
tausmag	MAG OF MOMENTUM SURFACE DRAG	[lon][lat]
tausmag_hemis	tausmag_hemis	—
tauus	U COMPON OF MOMENTUM SRF DRAG	[lon][lat]
tauus_hemis	tauus_hemis	—
tauvs	V COMPON OF MOMENTUM SRF DRAG	[lon][lat]
tauvs_hemis	tauvs_hemis	—
TEMPSI	SEA ICE TEMPERATURE (MASS LAYER 2)	[lon][lat]

**Example:
Screenshot
of NASA
GISS
model
output**

How to use a model



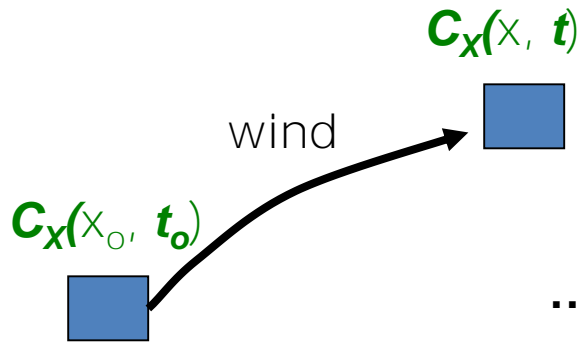
Boeing Delta II Rocket



Artist's rendering of Aura in Orbit

Need to be in constant dialogue.

Lagrangian model: Follow air parcel moving with wind

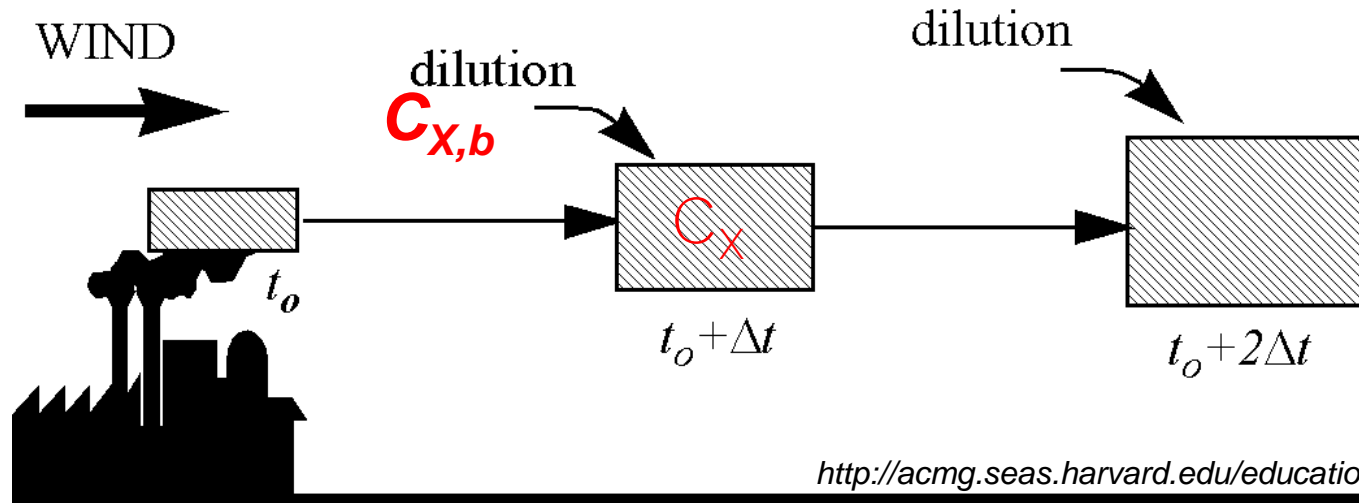


In the moving box (\mathbf{C} is the concentration),

$$\frac{dC_X}{dt} = E + P - L - D$$

...no transport terms! (they're implicit in the trajectory)

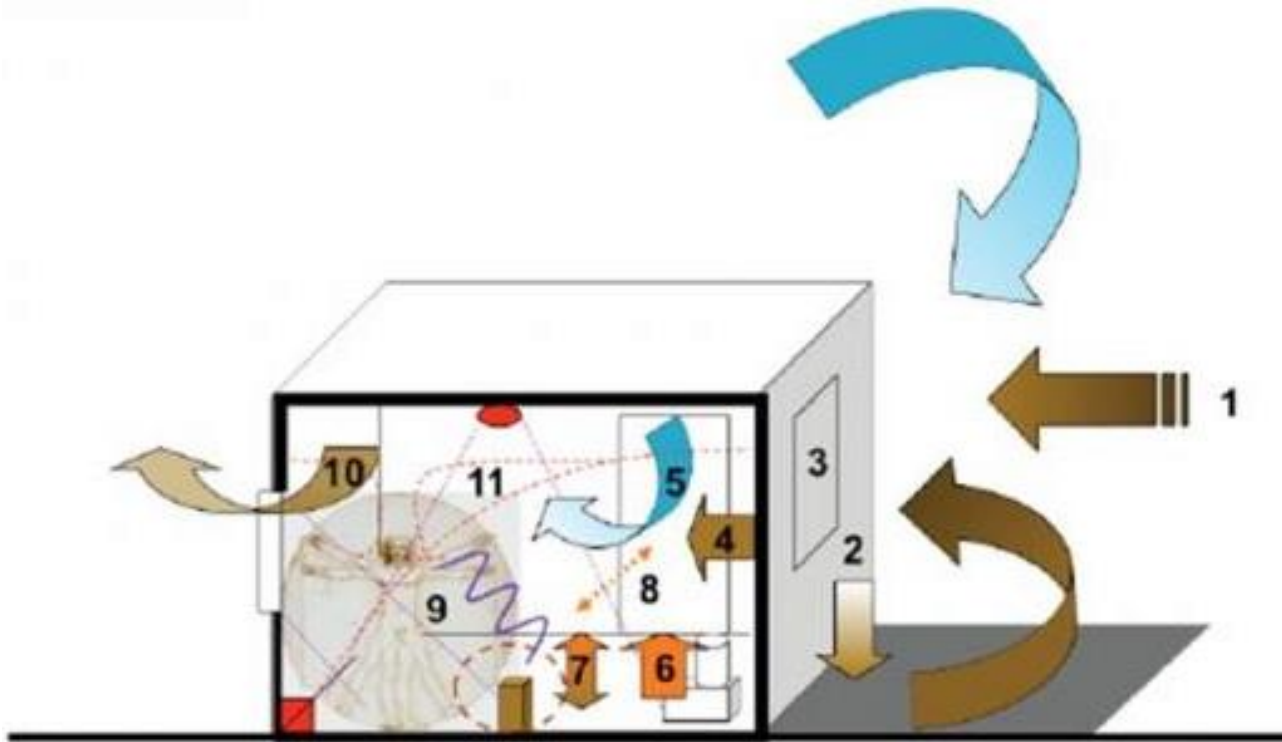
Application to the chemical evolution of an isolated pollution plume:



<http://acmg.seas.harvard.edu/education/>

In pollution plume,
$$\frac{dC_X}{dt} = E + P - L - D - k_{dilution} (C_X - C_{X,b})$$

Even indoor air pollution models!



- In many ways similar to what we saw earlier.

- Atmospheric models share several of their principles/characteristics, from the (very) **local** to the **global** scale.

1. Levels of outdoor-PM and its distributions
2. Penetration loss of outdoor-PM concentrations
3. Building characteristics
4. Outdoor-PM indoors gain through natural ventilation and infiltration
5. Outdoor clean air through natural ventilation
6. Levels of indoor-PM from indoor sources
7. Deposition loss and re-suspension of PM
8. Indoor-PM concentrations loss and gain through other rooms and/or indoor passageways
9. Human/animal presence and activity
10. Indoor-PM loss through natural ventilation and exfiltration
11. Interaction of and between pollutants (i.e. formation, phase change, coagulation)

Summary

- Showed and discussed the **mass balance equation** for atmospheric constituents in a model.
- Discussed **one-box** and **two-box** models.
- Presented basic **thermodynamic** and **dynamic** equations in models, and mentioned their interactions.
- Gave a summary of how models are **typically used**.
- Briefly mentioned **Lagrangian** and **indoor** air quality models.